

Polarization Leakage and Calibration

Michael Johnson (CfA)

with Vincent Fish, Kazu Akiyama, Rusen Lu, Keiichi Asada, Thomas Krichbaum, Lindy Blackburn, Shep Doeleman, Dick Plambeck, Dan Marrone, John Wardle, and the Event Horizon Telescope Collaboration

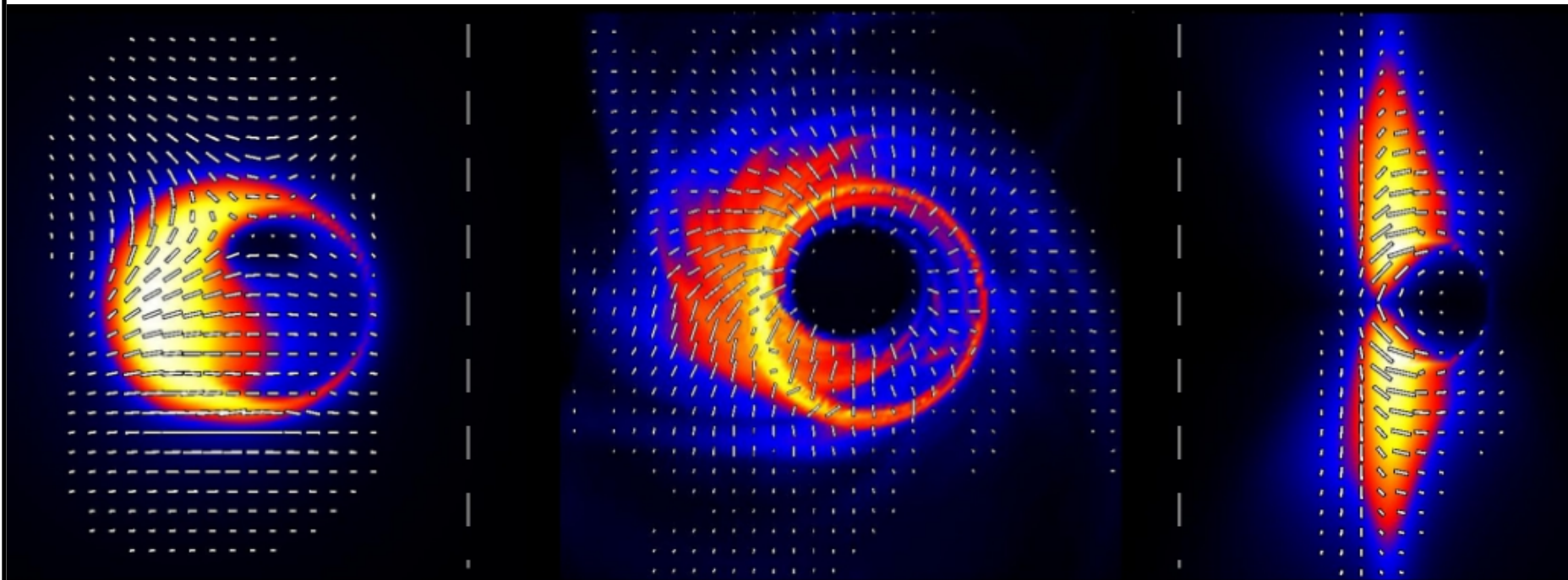
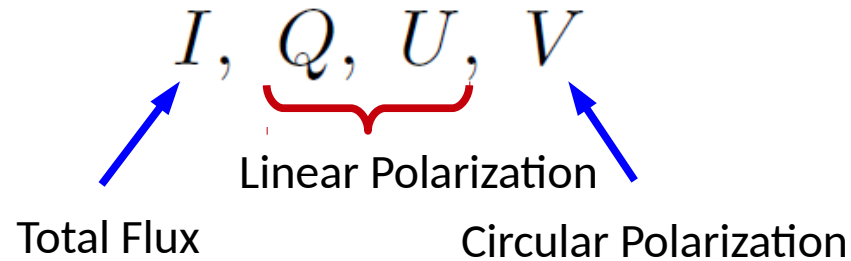
mm-VLBI Data Processing Workshop, Leiden

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Polarimetry with VLBI

Goal: Study images of each of the Stokes parameters



Simulated Images Courtesy of Avery Broderick and Jason Dexter

Polarimetry with VLBI

Coherency Matrix

$$\begin{pmatrix} \langle E_{1,R} E_{2,R}^* \rangle & \langle E_{1,R} E_{2,L}^* \rangle \\ \langle E_{1,L} E_{2,R}^* \rangle & \langle E_{1,L} E_{2,L}^* \rangle \end{pmatrix}$$

Stokes Parameters

$$\tilde{I}, \underbrace{\tilde{Q}, \tilde{U}}_{\text{Linear Polarization}}, \tilde{V}$$

Total Flux

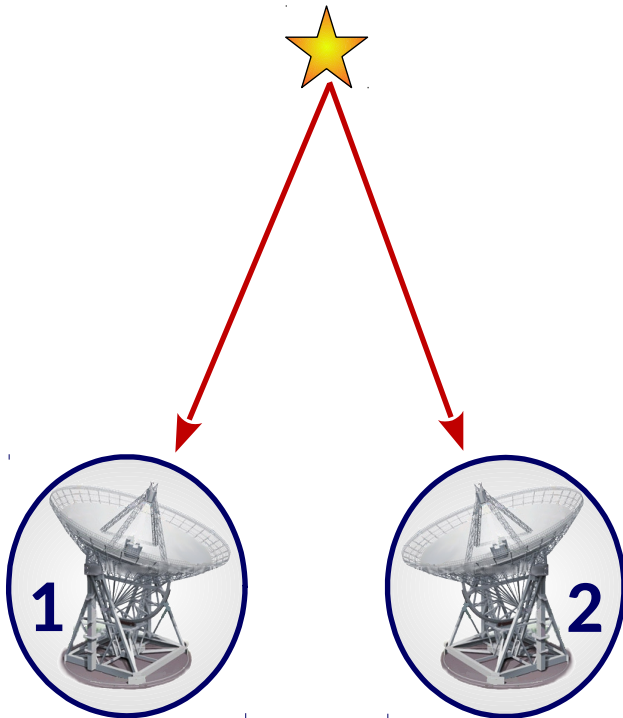
Circular Polarization

$$P = Q + iU$$

$$|Q|^2 + |U|^2 + |V|^2 \leq |I|^2$$

The phase of P (twice the **EVPA**) = polarization direction

Cross-hand fringes are usually much weaker!



Polarimetry with VLBI

Station actually measures:

$$\begin{pmatrix} E'_R \\ E'_L \end{pmatrix} = \underbrace{\begin{pmatrix} G_R & 0 \\ 0 & G_L \end{pmatrix}}_{\text{"Gain"}} \underbrace{\begin{pmatrix} 1 & D_R \\ D_L & 1 \end{pmatrix}}_{\text{"Leakage"}} \underbrace{\begin{pmatrix} e^{-i\phi} & 0 \\ 0 & e^{i\phi} \end{pmatrix}}_{\text{Field Rotation}} \begin{pmatrix} E_R \\ E_L \end{pmatrix} \equiv \mathbf{J} \begin{pmatrix} E_R e^{-i\phi} \\ E_L e^{i\phi} \end{pmatrix}$$

Jones Matrix
(Lorentz Transformation)

Leakage introduces spurious linear polarization, often exceeding the signal

The field rotation is critical – the source and instrumental contributions respond differently!

Notice: This description is arbitrary. Specifying conventions is essential.

Fractional Polarization

Intrinsic fractional polarization

$$\frac{R_1 L_2^*}{R_1 R_2^*} \approx \left(\frac{G_{2,L}}{G_{2,R}} \right)^* \left[m e^{-2i\phi_2} + \underbrace{D_{1,R} e^{2i\phi_{12}} + D_{2,L}^*}_{\text{"Leakage"}} \right]$$

↑ ↑

Unknown, but stable, phase

Gain amplitudes are easy with 3 sites:

$$\left| \frac{G_{1,R}}{G_{1,L}} \right| = \sqrt{\frac{\langle R_1 R_2^* \rangle}{\langle L_1 L_2^* \rangle} \times \frac{\langle R_3 R_1^* \rangle}{\langle L_3 L_1^* \rangle} \times \frac{\langle L_3 L_2^* \rangle}{\langle R_3 R_2^* \rangle}}$$

Idea: Phase reference (weak) cross-hand visibilities to (strong) parallel-hand visibilities

Immune to scatter broadening

Fourier relationship no longer applies!

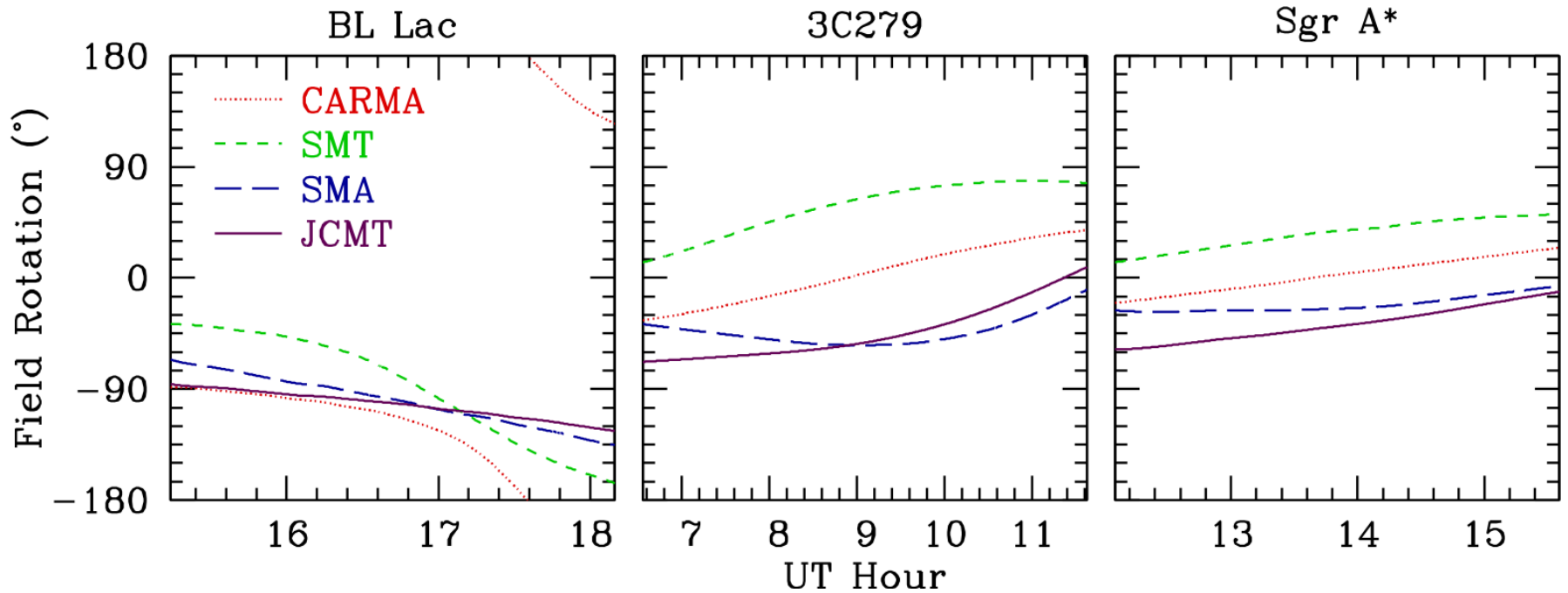
Calibration Degeneracies

$$\frac{R_1 L_2^*}{R_1 R_2^*} \approx \left(\frac{G_{2,L}}{G_{2,R}} \right)^* \left[m e^{-2i\phi_2} + \underbrace{D_{1,R} e^{2i\phi_{12}} + D_{2,L}^*}_{\text{"Leakage"}} \right]$$

Notice: For sites with identical field rotation angles ϕ_x , D_R and D_L don't decouple

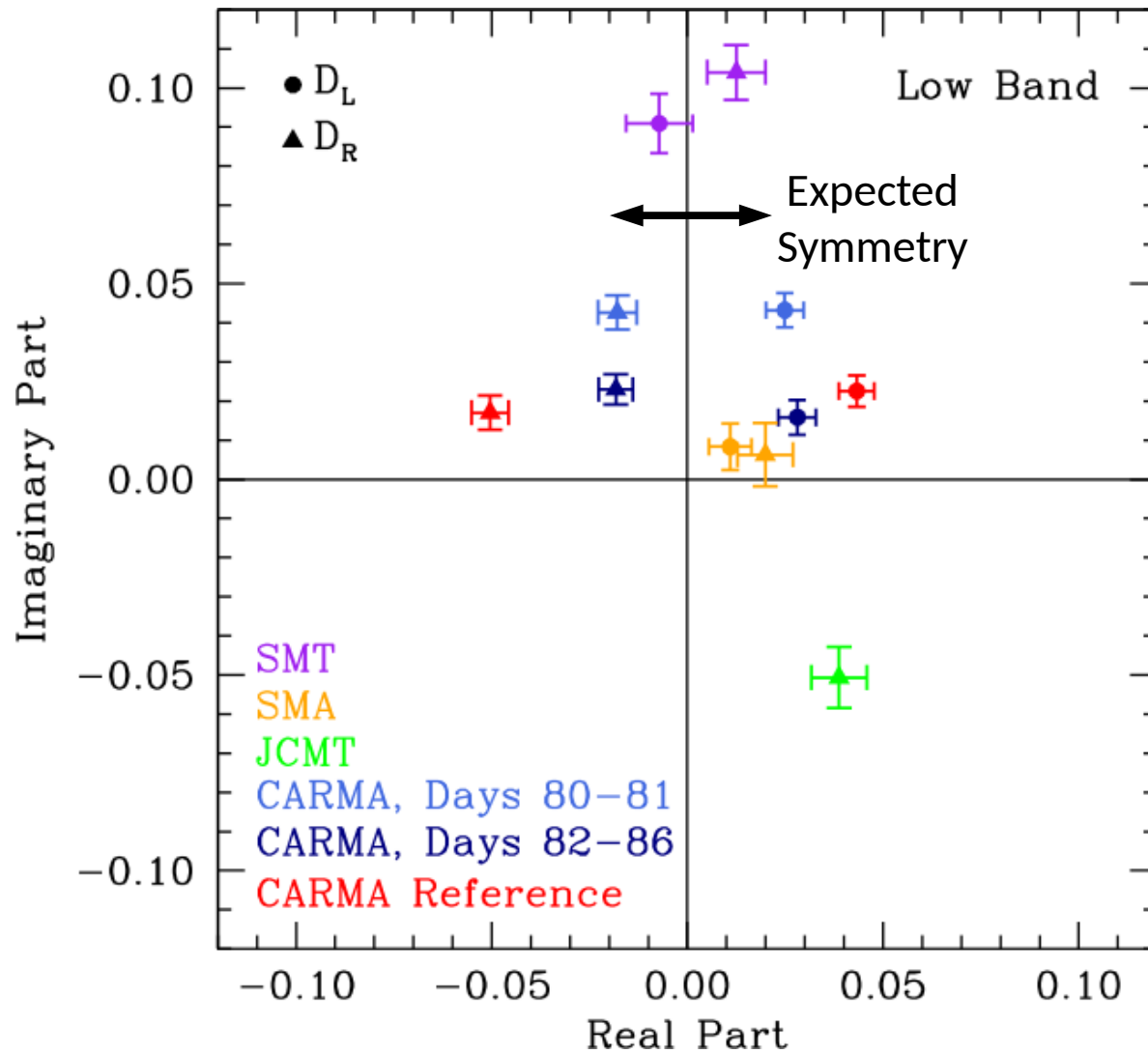
This is the case with most arrays (e.g., the SMA and CARMA)

Field Rotation Angles

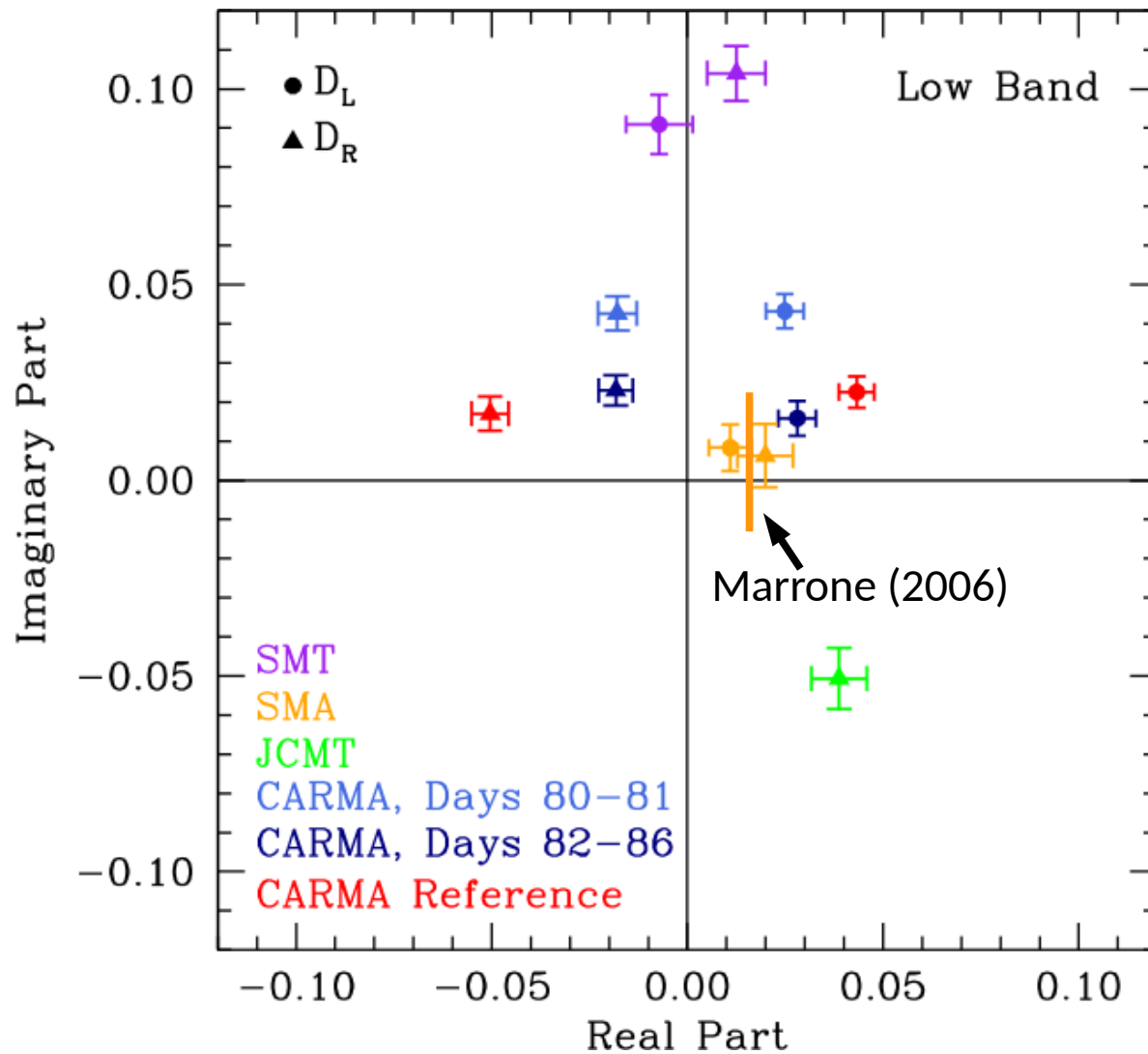


Remember, leakage terms rotate by twice the field-rotation angle

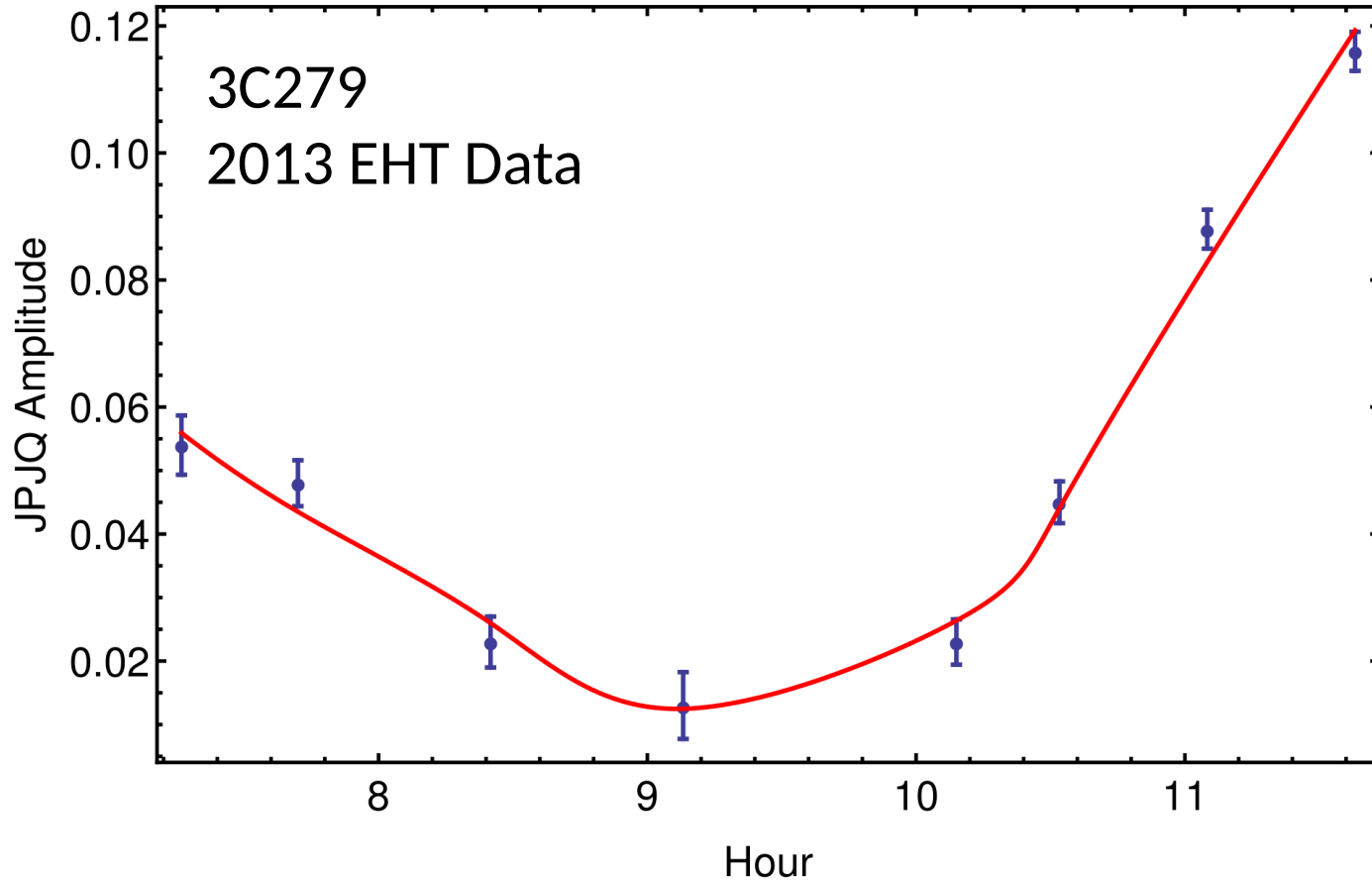
Derived Leakage Terms



Leakage Terms vs. Array Measurements



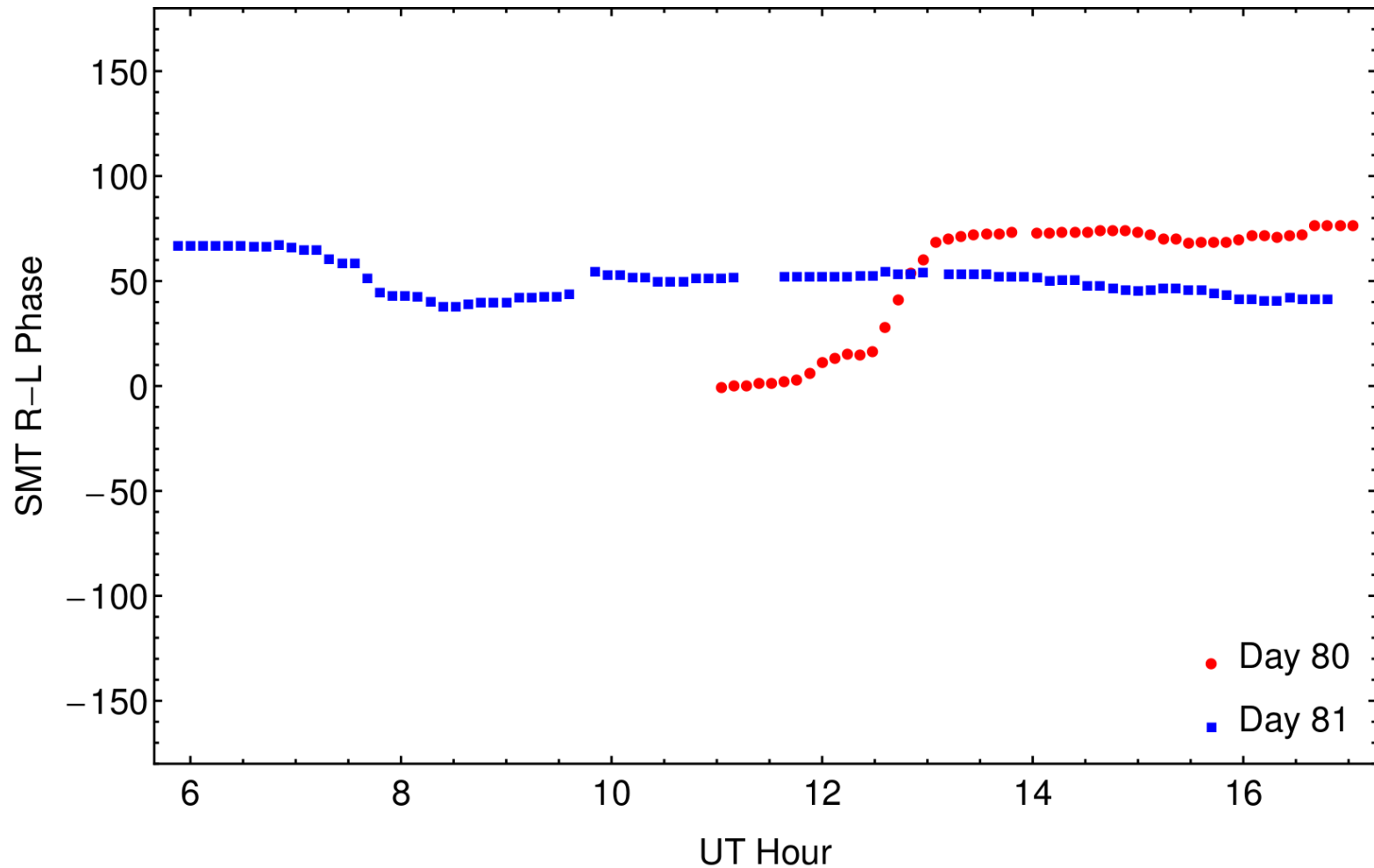
JCMT-SMA



Note: The SMA and the JCMT have different field rotation angles

Leakage effects show up in fractional polarization amplitudes for non-simultaneous scans!

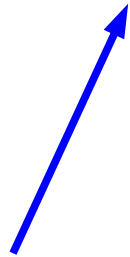
SMT R/L Phase Drift



Unexpected! We need to be careful about building a framework upon false assumptions.

Leakage can Affect Visibility Amplitudes!

$$\langle R_1 R_2^* \rangle \approx \tilde{I}(\mathbf{u}_{12}) G_{1,R} G_{2,R}^* e^{-i(\phi_1 - \phi_2)} [1 + \check{m}^*(-\mathbf{u}_{12}) D_{1,R} e^{2i\phi_1} + \check{m}(+\mathbf{u}_{12}) D_{2,R}^* e^{-2i\phi_2} + \check{v}(\mathbf{u}_{12})]$$
$$\langle L_1 L_2^* \rangle \approx \tilde{I}(\mathbf{u}_{12}) G_{1,L} G_{2,L}^* e^{+i(\phi_1 - \phi_2)} [1 + \check{m}(+\mathbf{u}_{12}) D_{1,L} e^{-2i\phi_1} + \check{m}^*(-\mathbf{u}_{12}) D_{2,L}^* e^{2i\phi_2} - \check{v}(\mathbf{u}_{12})]$$



Up to ~10% effect on amplitudes
Introduces non-closing phase errors!

Summary

Polarization is a rich source of information for VLBI but precise calibration is essential.

Polarization leakage affects many things

- Spurious linear polarization
- Biased visibility amplitudes
- Biased closure phases

Calibration for past EHT data appears to be secure

- Leakage terms are $\lesssim 10\%$
- No evidence yet for strong elevation-dependent effects
- The role of circular polarization isn't yet clear

The most difficult problems to calibrate are those we don't expect

- SMT R-L phase drift

Summary

But... we're learning that the EHT can be very well calibrated

Differences from other arrays:

1. Co-located sites with different field rotation angles (SMT and JCMT)
2. Strong polarization on long baselines
3. Heterogeneous array
4. Long observing tracks

Calibration corrections are large but accurate