## EHT Data: Overview and Challenges

Vincent L. Fish

#### Observe

Correlate

Reduce

Calibrate

#### Observe

Test and operate equipment, scheduling, telescope control, monitoring, logs, ...

Correlate

#### Reduce

Calibrate

#### Observe

Test and operate equipment, scheduling, telescope control, monitoring, logs, ...

#### Correlate

Theoretically simple, but substantial bookkeeping overhead and numerous ways for things to go wrong

### Reduce

Calibrate

#### Observe

Test and operate equipment, scheduling, telescope control, monitoring, logs, ...

#### Correlate

Theoretically simple, but substantial bookkeeping overhead and numerous ways for things to go wrong

### Reduce

Find residual quantities (delay, rate, ...) relative to correlator model, estimate correlation coefficients and produce other observables

### Calibrate

#### Observe

Test and operate equipment, scheduling, telescope control, monitoring, logs, ...

#### Correlate

Theoretically simple, but substantial bookkeeping overhead and numerous ways for things to go wrong

### Reduce

Find residual quantities (delay, rate, ...) relative to correlator model, estimate correlation coefficients and produce other observables

### Calibrate

Incorporate metadata, obtain self-consistent gain solutions, determine polarization leakages, and generate final data products

#### Observe

Test and operate equipment, scheduling, telescope control, monitoring, logs, ...

#### Correlate

Theoretically simple, but substantial bookkeeping overhead and numerous ways for things to go wrong

### Reduce

Find residual quantities (delay, rate, ...) relative to correlator model, estimate correlation coefficients and produce other observables

#### Calibrate

Incorporate metadata, obtain self-consistent gain solutions, determine polarization leakages, and generate final data products

#### Analyze and interpret

Fit models to data or produce image, compare with simulations Generate science Write papers

# Challenge of high-frequency VLBI

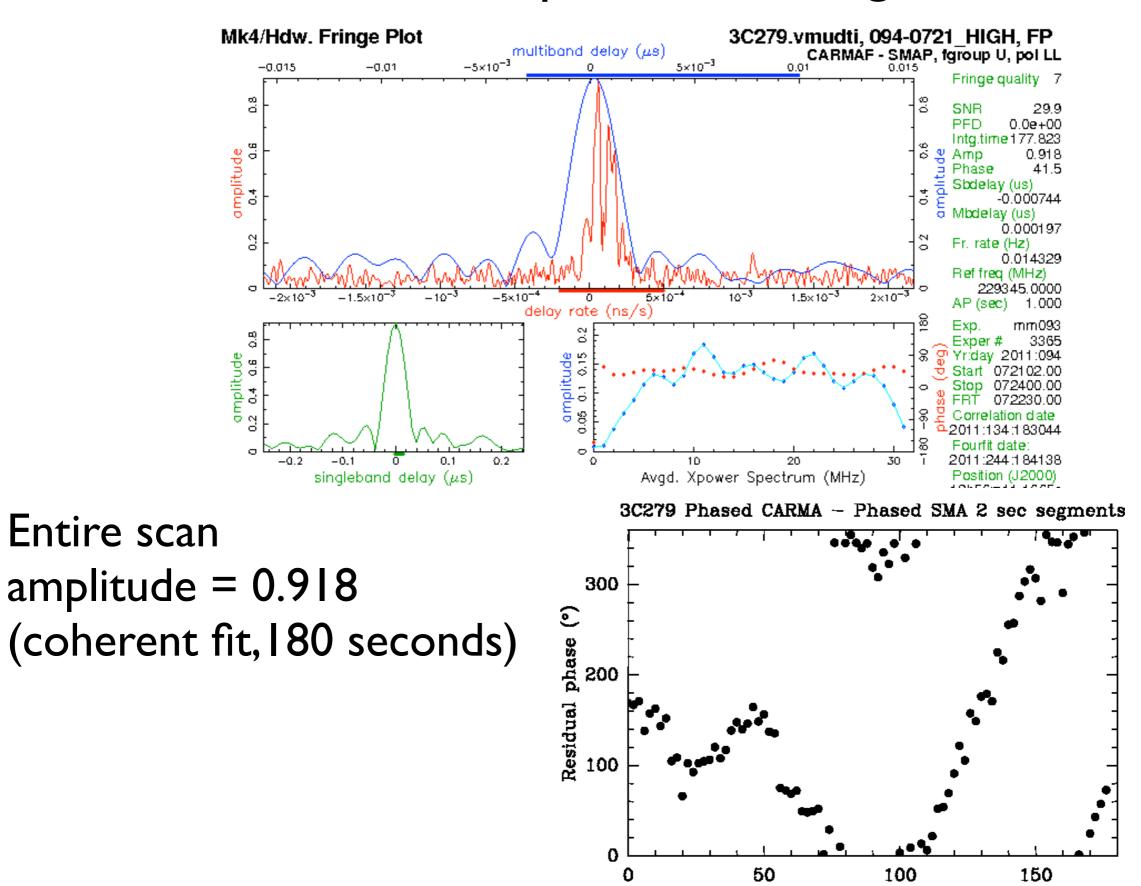
In one word: Atmosphere

Two main issues Opacity Coherence

Rapidly changing tropospheric delays introduce phase variations

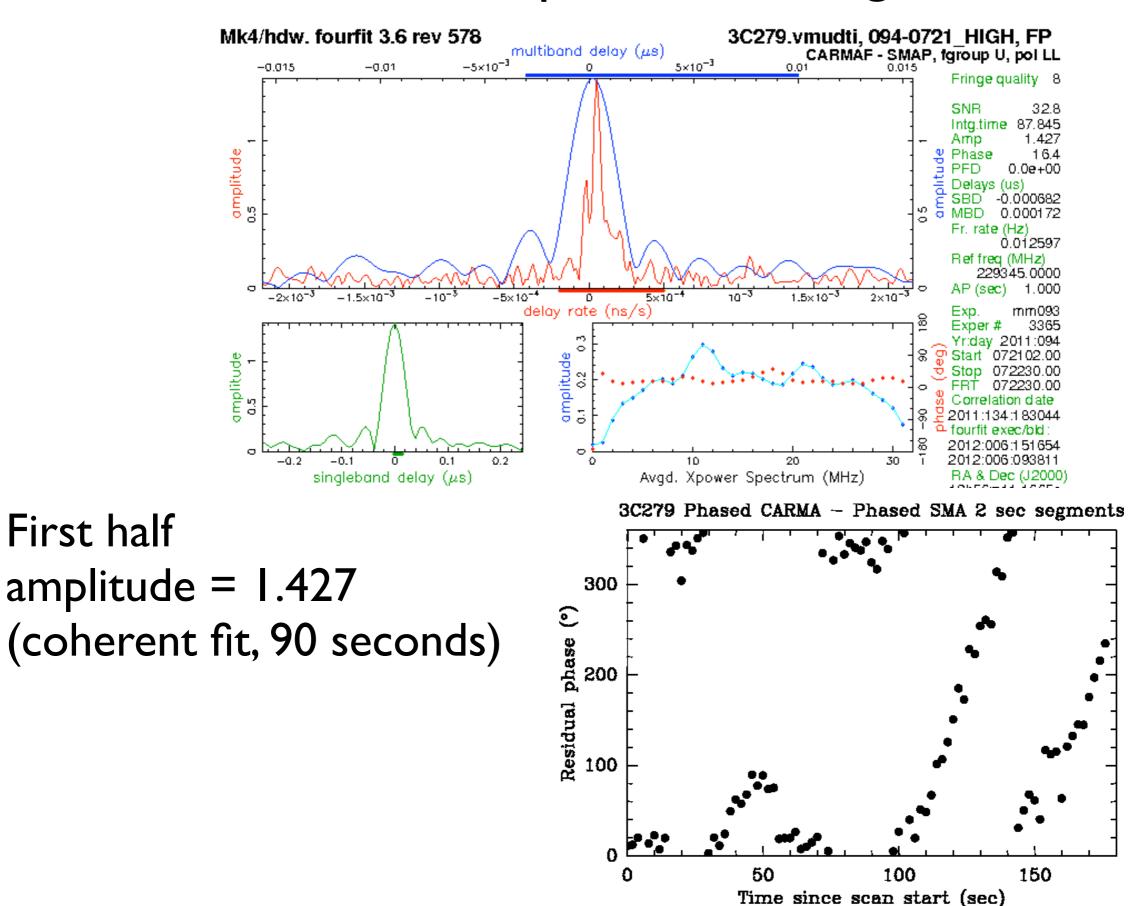
Effects on the data Sensitivity loss Amplitude loss Corruption of phase information

### Example scan: Strong

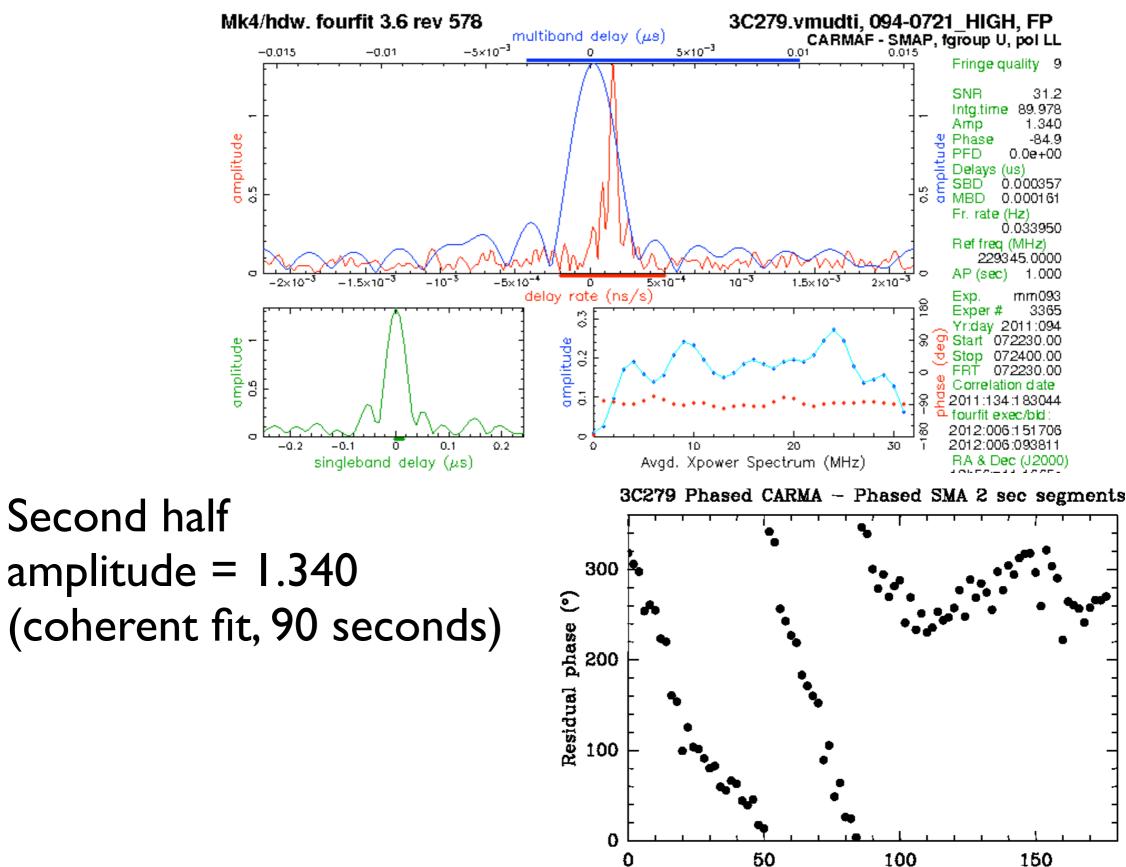


Time since scan start (sec)

### Example scan: Strong



### Example scan: Strong

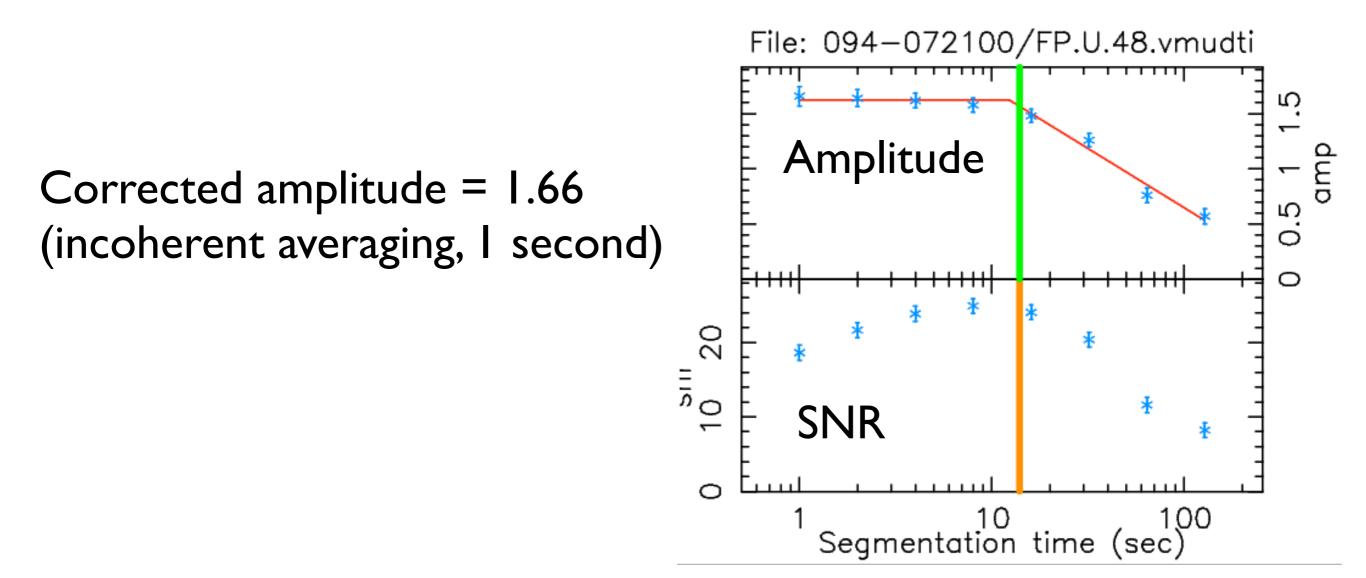


Time since scan start (sec)

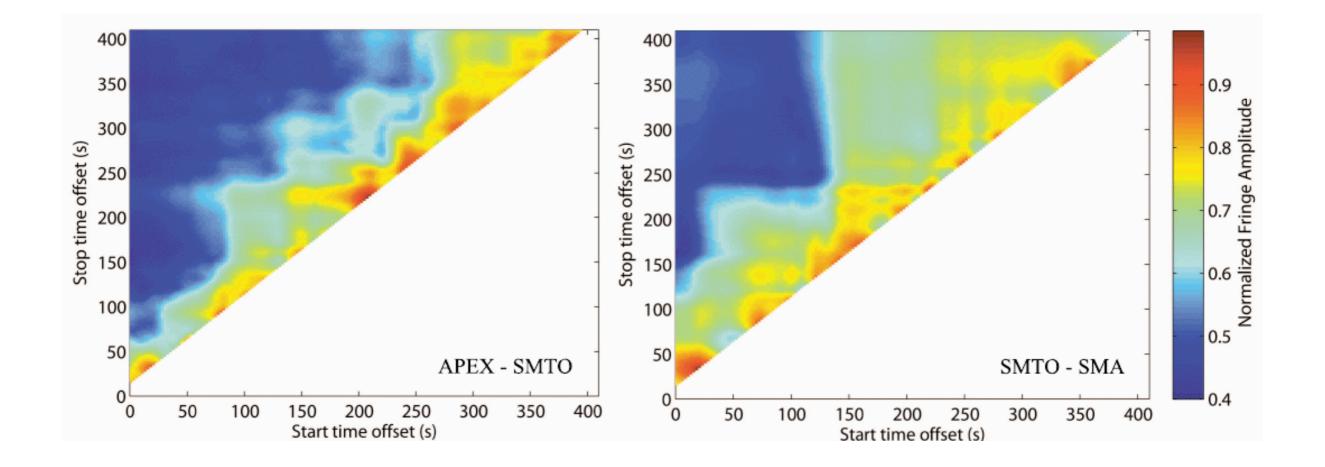
## Atmosphere and coherence

The atmosphere introduces rapid phase fluctuations

Must segment data and average incoherently to optimize detection as well as to determine the correlation coefficient



Not easy!



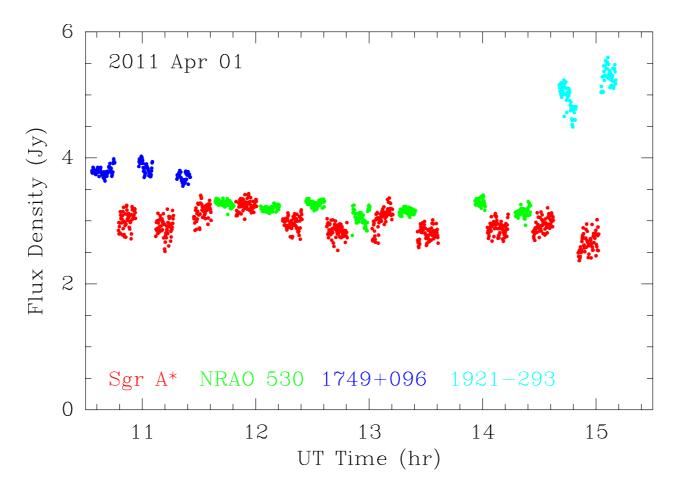
#### Not easy!

3C273-LL M87-LL3C279-LL• 1633+382-LL 3C345-LL 1749+096-LL NRA0530-LL 10MWC349-LL 1921-293-LL Amplitude BLLAC-LL ß 8<sup>h</sup>  $10^{h}$ 16<sup>h</sup> 6<sup>h</sup> 12<sup>h</sup>  $14^{h}$ 

LL 222.3210 GHz 1.00m

Time (courtesy D. Plambeck)

Not easy!



(courtesy D. Plambeck)

#### Not easy!

3C273-LL M87-LL3C279-LL• 1633+382-LL 3C345-LL 1749+096-LL NRA0530-LL 10MWC349-LL 1921-293-LL Amplitude BLLAC-LL ß 8<sup>h</sup>  $10^{h}$ 16<sup>h</sup> 6<sup>h</sup> 12<sup>h</sup>  $14^{h}$ 

LL 222.3210 GHz 1.00m

Time (courtesy D. Plambeck)

#### Not easy!

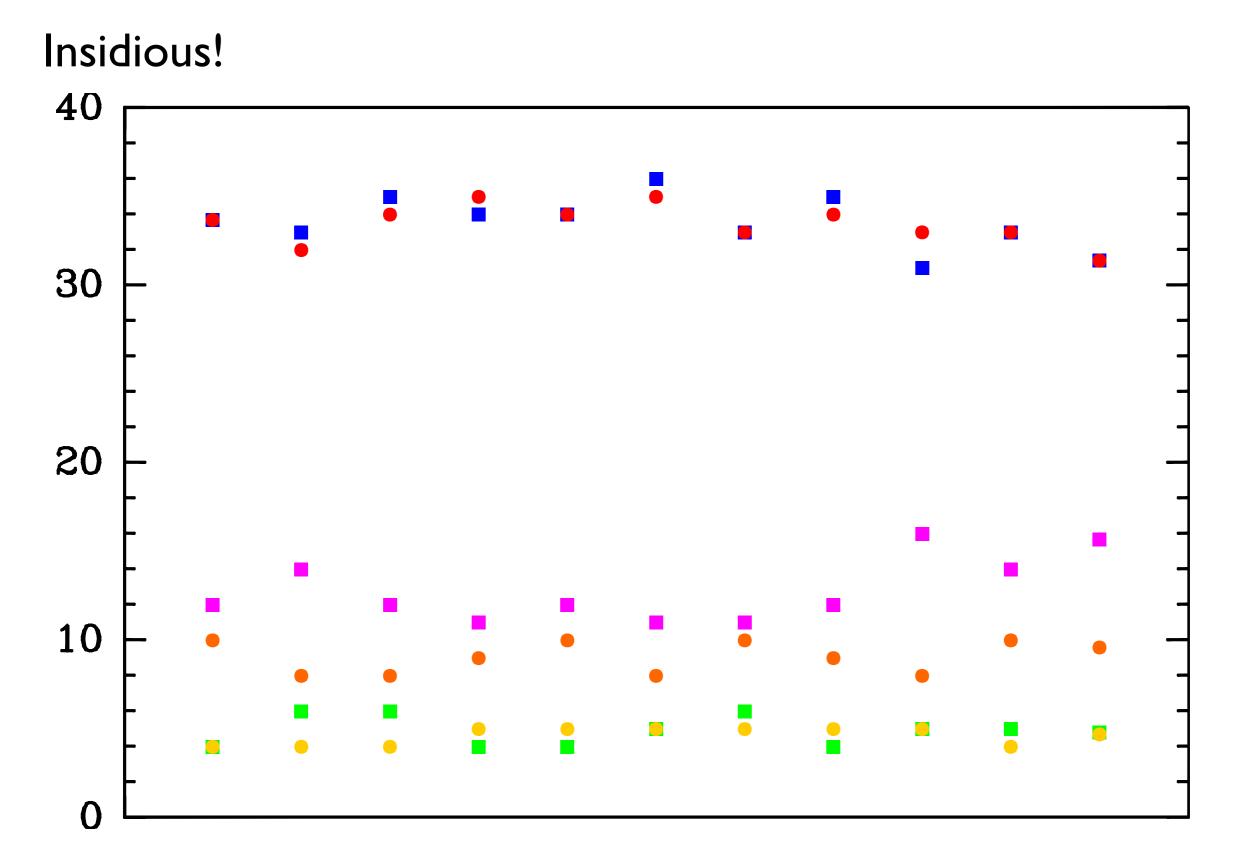
3C273-LL M87-LL15 3C279-LL 1749+096-LL 1633+382-LL • SGRA-LL MWC349-LL 1921-293-LL 10 BLLAC-LL Amplitude ß  $8^{h}$  $10^{h}$  $12^{h}$  $14^{h}$ 16<sup>h</sup>  $6^{\rm h}$ Time (courtesy D. Plambeck)

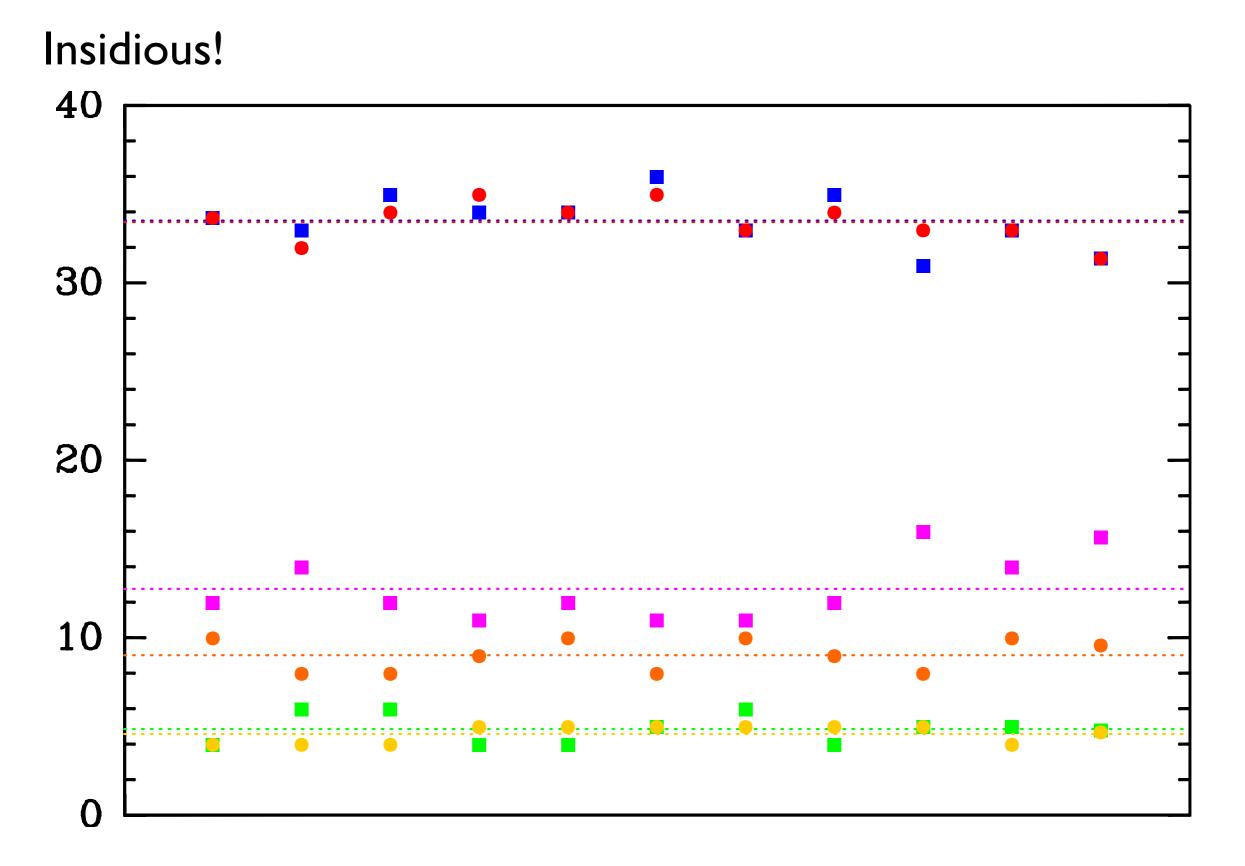
LL 222.3210 GHz 1.00m

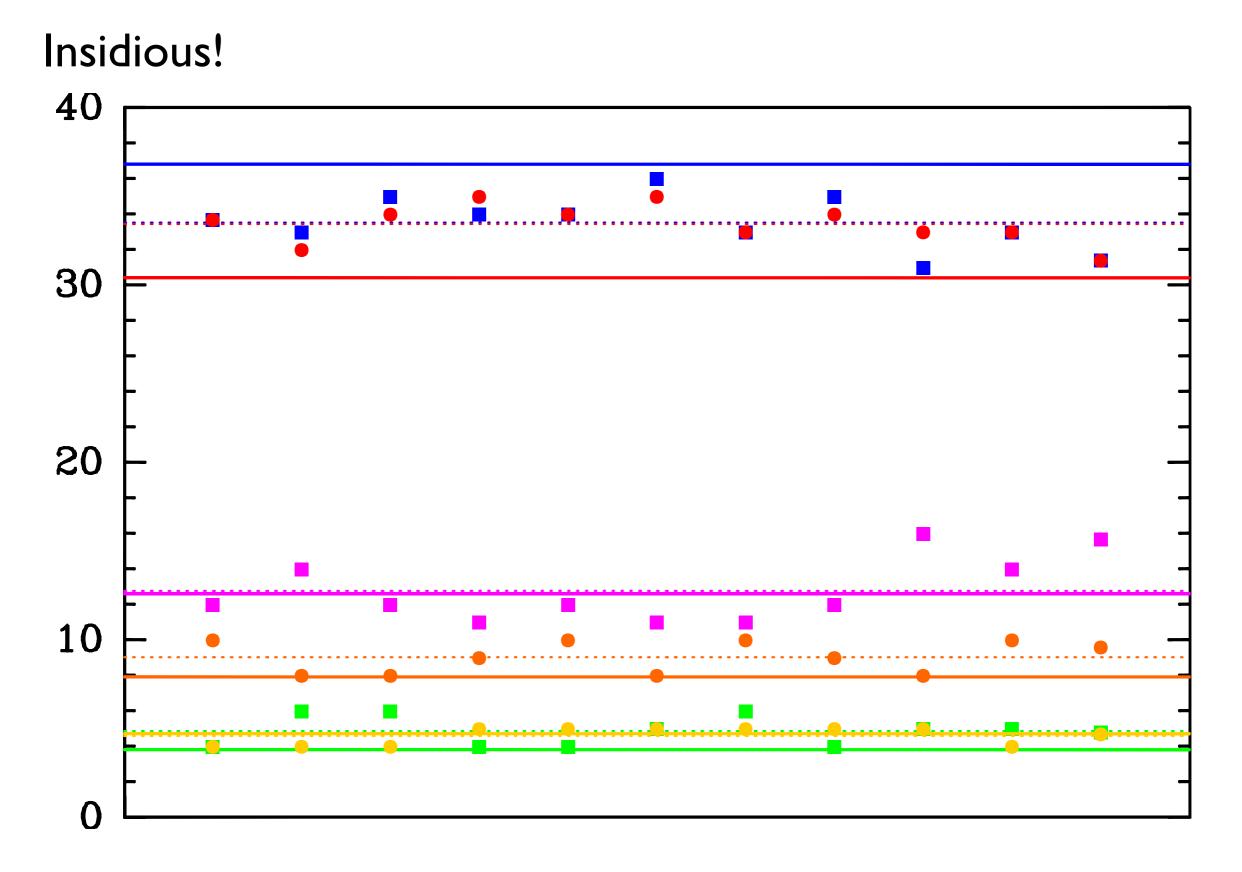
Not easy!

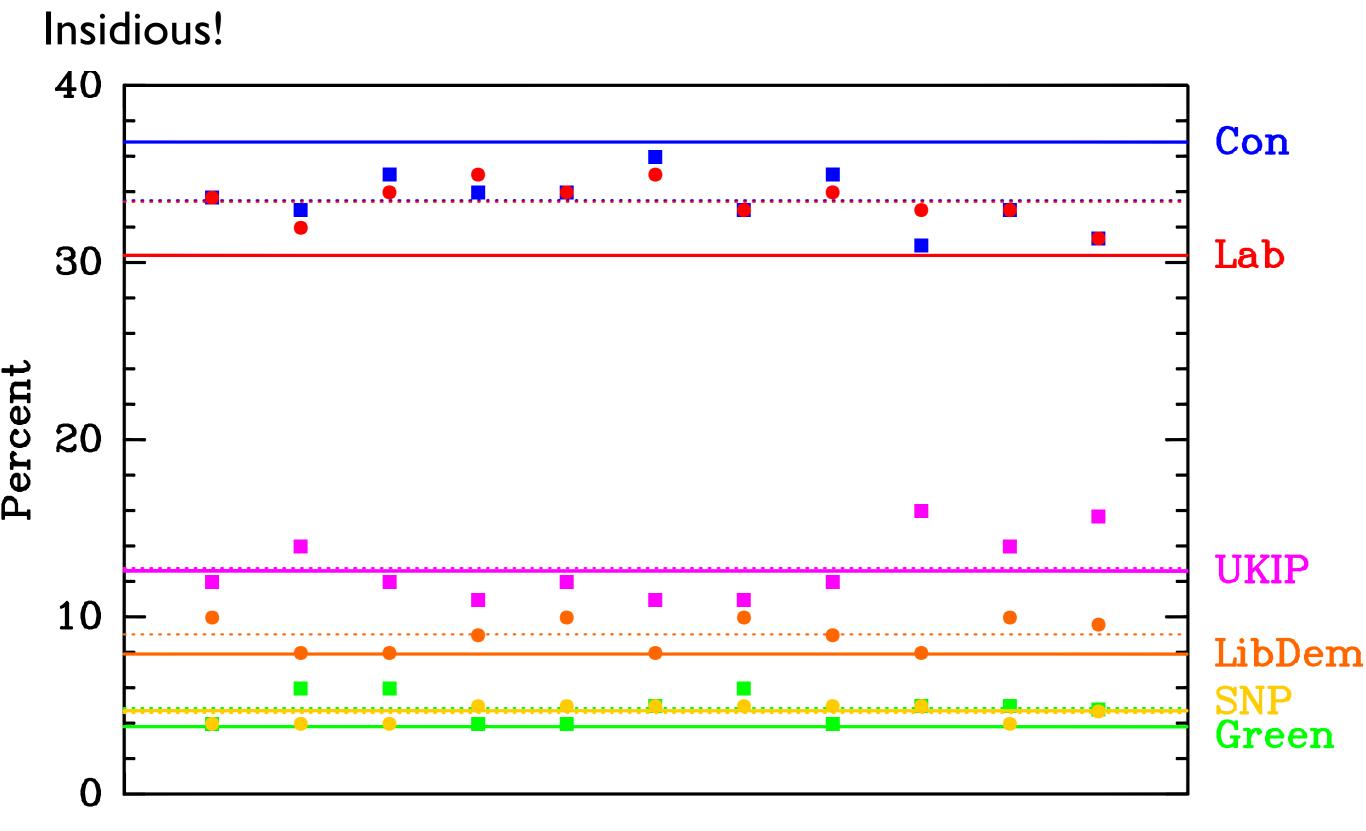
Hard for connected-element and single-dish measurements too

Potential difficulties at several stages Determining correlation coefficients Getting high-quality calibration information (T<sub>sys</sub>, tau, gain, ...) Measurement errors in these quantities too Infrequent measurements Directionality Systematic errors









Final polls of 2015 UK election

Insidious!

Often neglected entirely

Treating systematic errors as random errors leads to high confidence in an erroneous result

Interpretation of data and comparison with models/simulations should take into account (possibly unknown) systematic errors

<u>What it is</u>

Specialized toolkit for mmVLBI processing Highly flexible fringe finding with diagnostic plots

<u>What it is</u>

Specialized toolkit for mm VLBI processing

Highly flexible fringe finding with diagnostic plots

Coherent and incoherent searching

<u>What it is</u>

Specialized toolkit for mm VLBI processing Highly flexible fringe finding with diagnostic plots Coherent and incoherent searching Amplitude estimation with noise debiasing

<u>What it is</u>

Specialized toolkit for mm VLBI processing Highly flexible fringe finding with diagnostic plots Coherent and incoherent searching Amplitude estimation with noise debiasing Text files for easy input/output and human readability

<u>What it is</u>

Specialized toolkit for mm VLBI processing Highly flexible fringe finding with diagnostic plots Coherent and incoherent searching Amplitude estimation with noise debiasing Text files for easy input/output and human readability

<u>What it isn't</u>

An end-to-end solution: calibration and imaging done externally

<u>What it is</u>

Specialized toolkit for mm VLBI processing Highly flexible fringe finding with diagnostic plots Coherent and incoherent searching Amplitude estimation with noise debiasing Text files for easy input/output and human readability

<u>What it isn't</u>

An end-to-end solution: calibration and imaging done externally

More on HOPS in the tutorial