

Status of GGAO and Westford VGOS Stations

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Overview

- VGOS in one slide
- Hardware status
 - Overall signal chain
 - Feeds / frontends
- Recent measurements
- Future enhancements
 - Field System support
 - Wider up/down converter frequency range
 - Cable delay measurement system

VGOS = VLBI Geodetic Observing System

- Primary technical goal
 - 1-mm position accuracy in 24 hours
- Limiting error sources
 - Variable atmospheric delay
 - Sensitivity
- Strategy
 - Use fast-slewing ($5^\circ/\text{s}$ - $12^\circ/\text{s}$) antennas
 - Reduce delay error through high data rate and wide spanned BW (broadband delay)
- Implementation specifics
 - Antennas of $\geq 12\text{m}$ diameter
 - Data rates ≥ 8 Gbps using 4 dual-pol bands of 0.5-1.0 GHz BW each
 - Bands spaced over 2.2 GHz to ~ 14 GHz

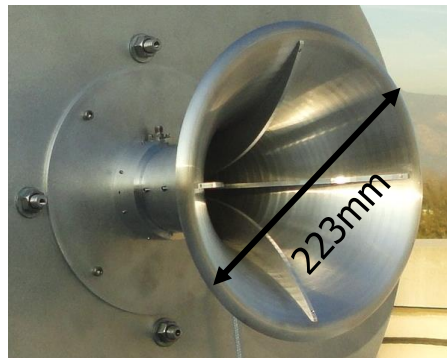
VGOS Signal Chain at GGAO and Westford

- Cryogenically cooled broadband QRFH (Quadruple-Ridged Flared Horn) feed
 - 2 linear-polarization coaxial outputs
 - Designed by Ahmed Akgiray (Caltech)
- Cryogenically cooled LNAs (Caltech CRYO1-12), 1 per polarization
- Pulse cal (5 MHz rep rate) and noise cal injection between feed and LNAs
- RF separated into low (coax or fiber downlink) and high (fiber) frequency bands
- Four up/down converters (UDCs)
- Four RDBE-H backends with 3.0 firmware
 - 1024 MHz sample clock
 - Two 512-MHz inputs (two polarizations)
 - Complex samples
 - Output = sixteen 32-MHz-BW PFB channels (8 dual-pol or 16 single-pol)
 - Pulse cal tone extraction
 - 10-100 Hz noise cal synchronous detection
- One Mark 6 recorder

GGAO antenna/feed



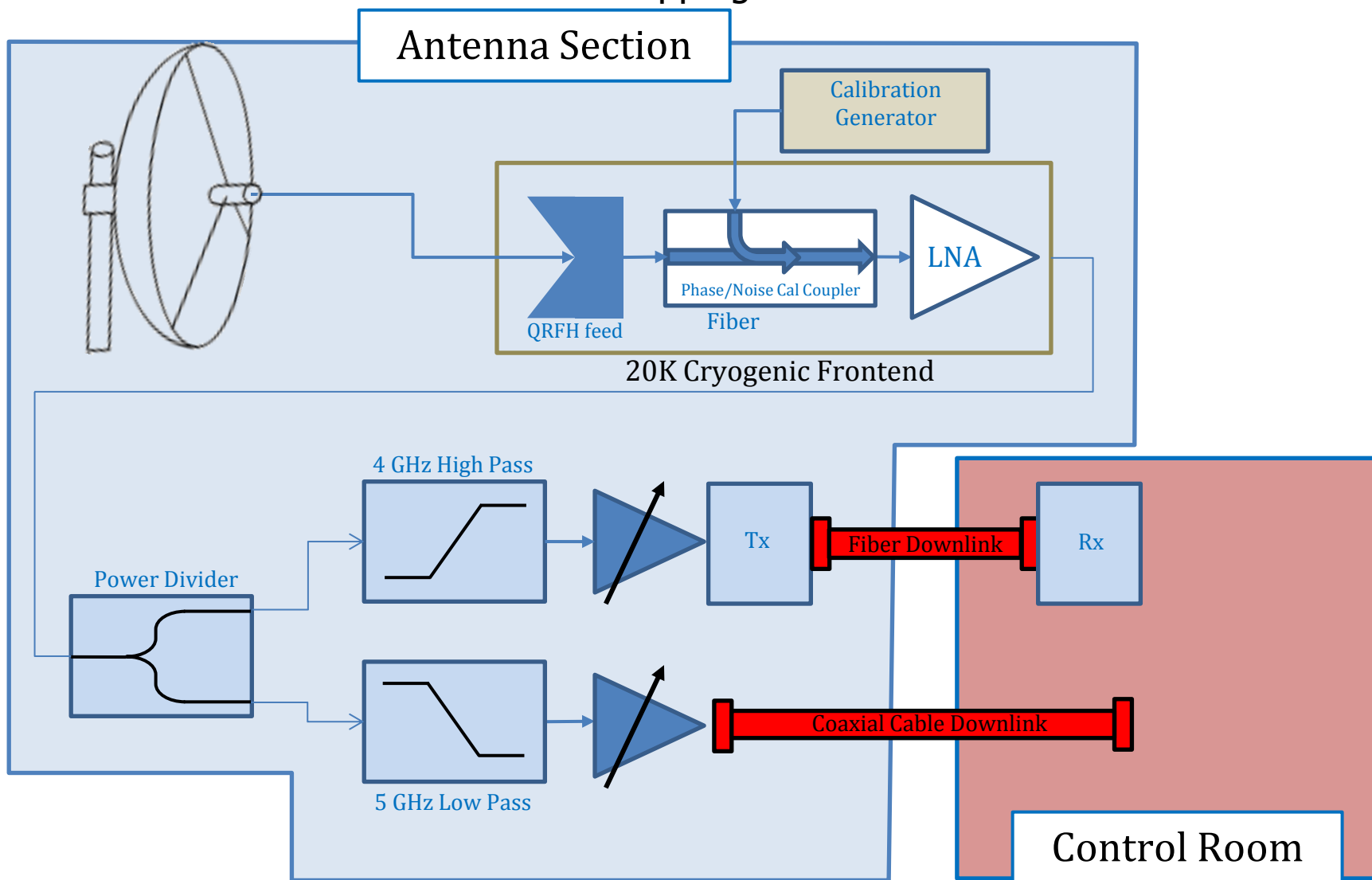
- 12 m diameter antenna
 - Cassegrain optics
 - Shaped reflectors
 - 50° feed optical half angle



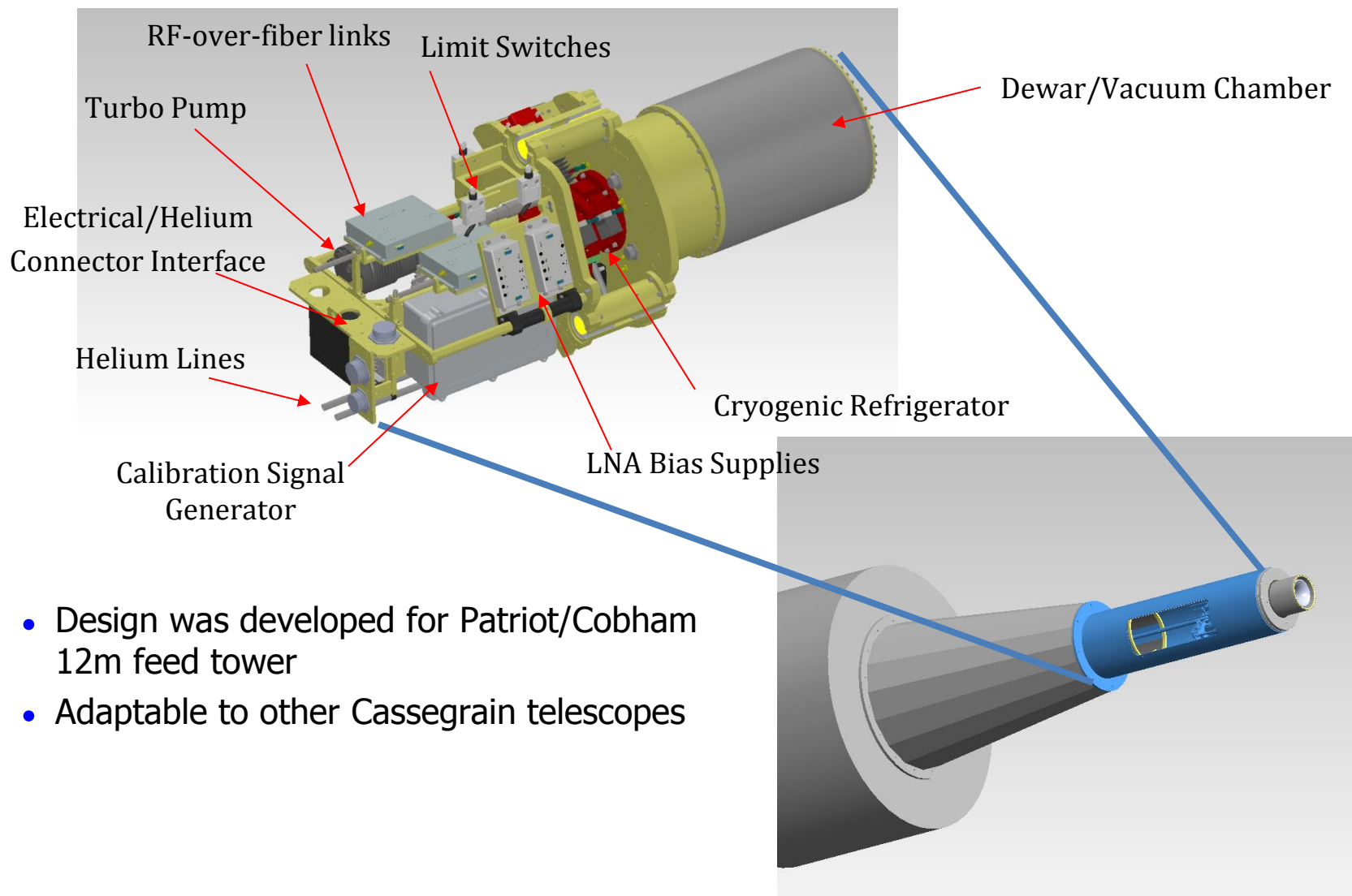
- QRFH feed
 - Optimized for
 - 12m optics
 - 2-12 GHz

GGAO receiver design

2 overlapping bands

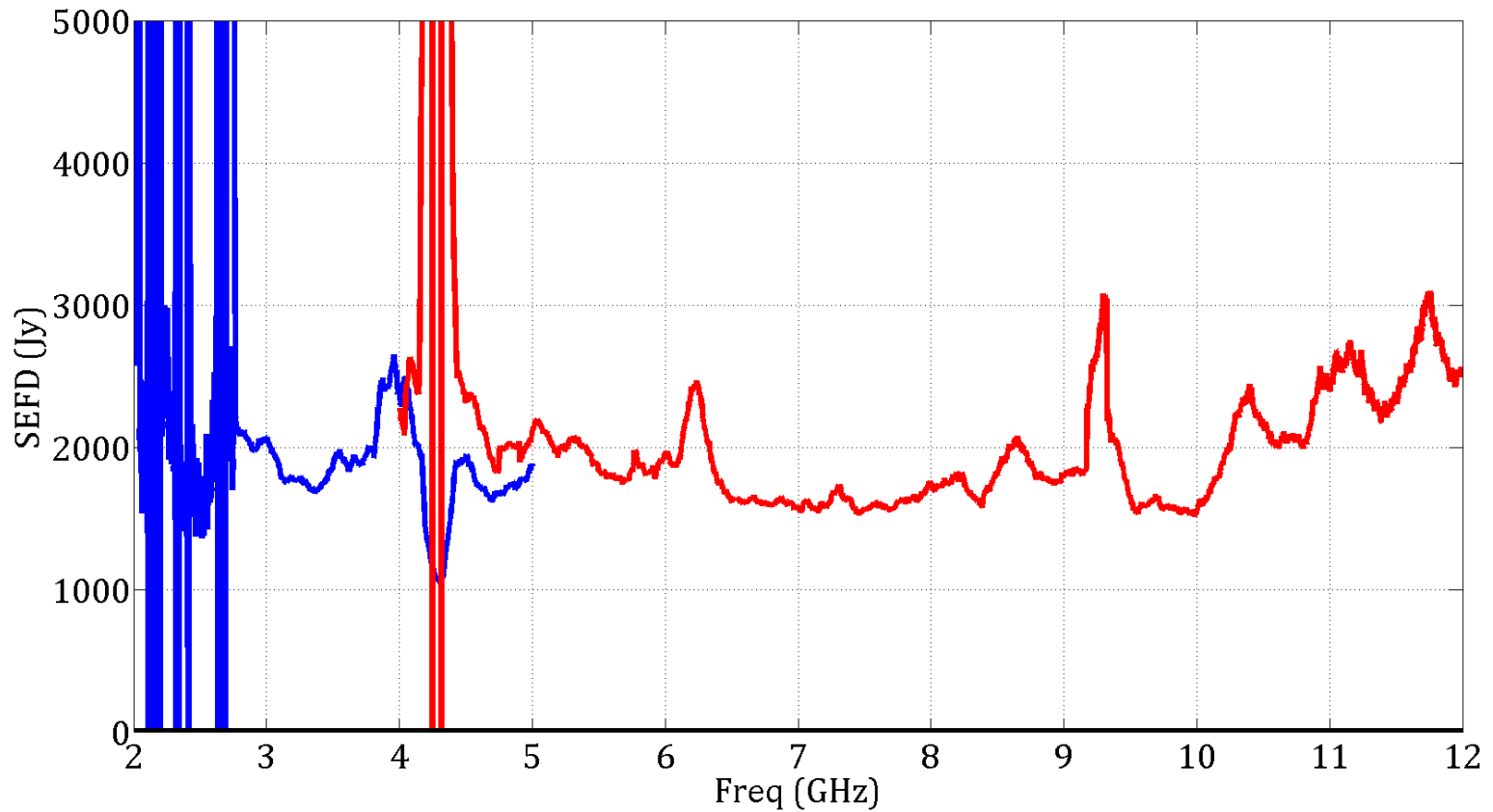


GGAO frontend mechanical design



- Design was developed for Patriot/Cobham 12m feed tower
- Adaptable to other Cassegrain telescopes

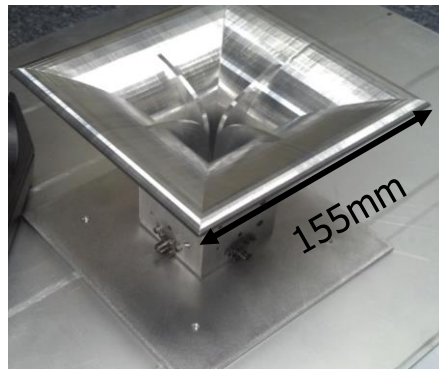
GGAO SEFD



Westford antenna/feed



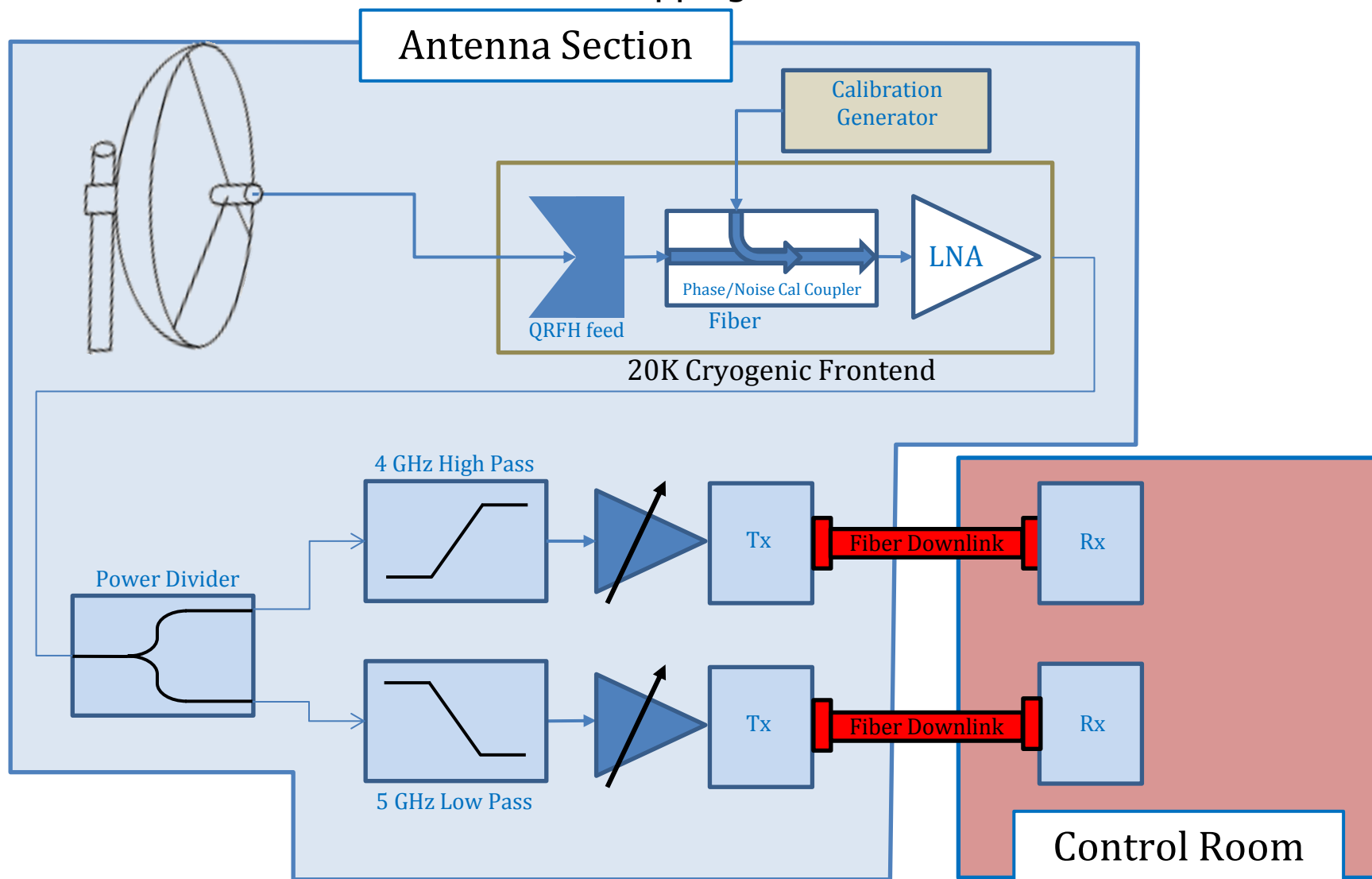
- 18 m diameter antenna
 - VGOS feed at prime focus
 - f/D 0.3 paraboloid primary
 - 80° feed optical half angle



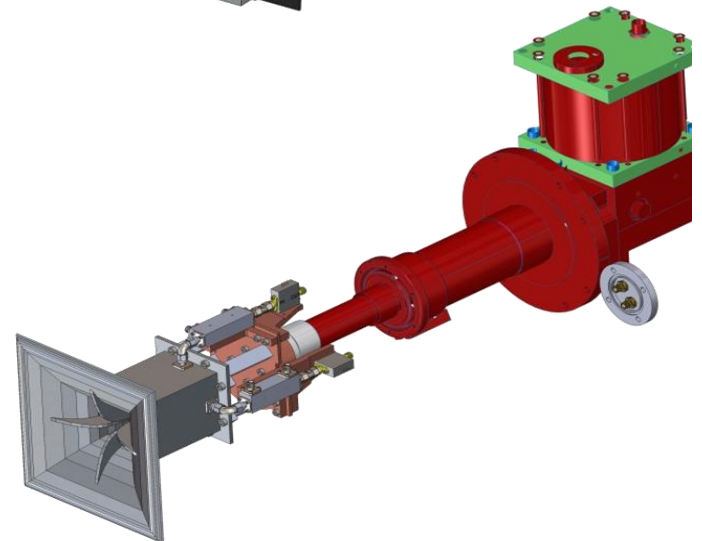
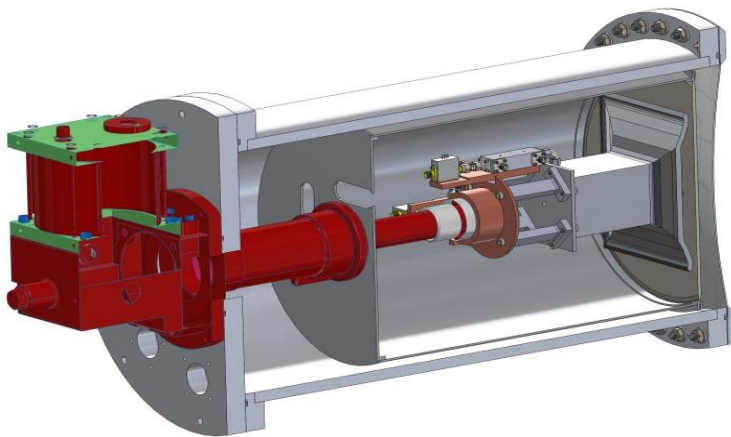
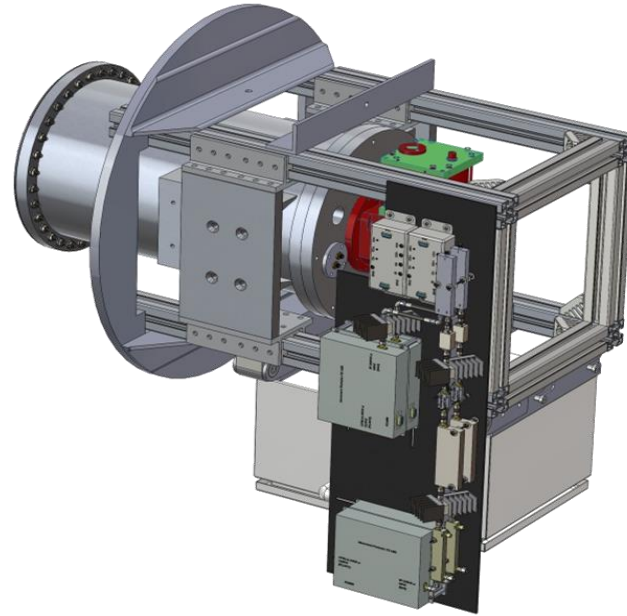
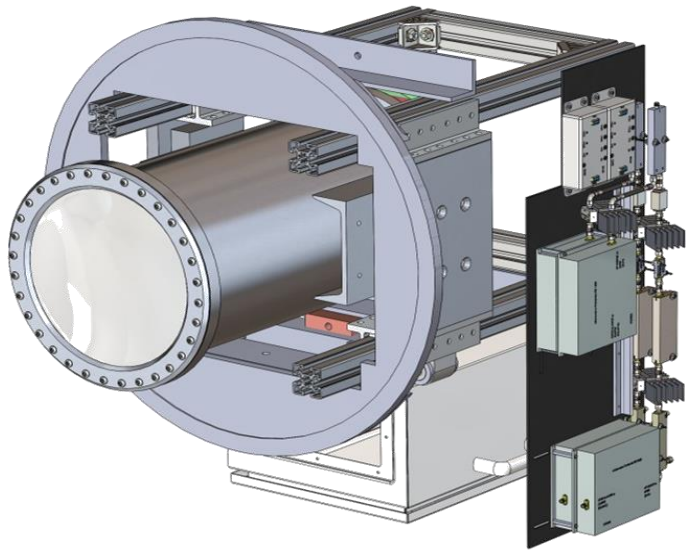
- QRFH feed
 - Optimized for
 - Westford prime-focus optics
 - 2-14 GHz

Westford receiver design

2 overlapping bands

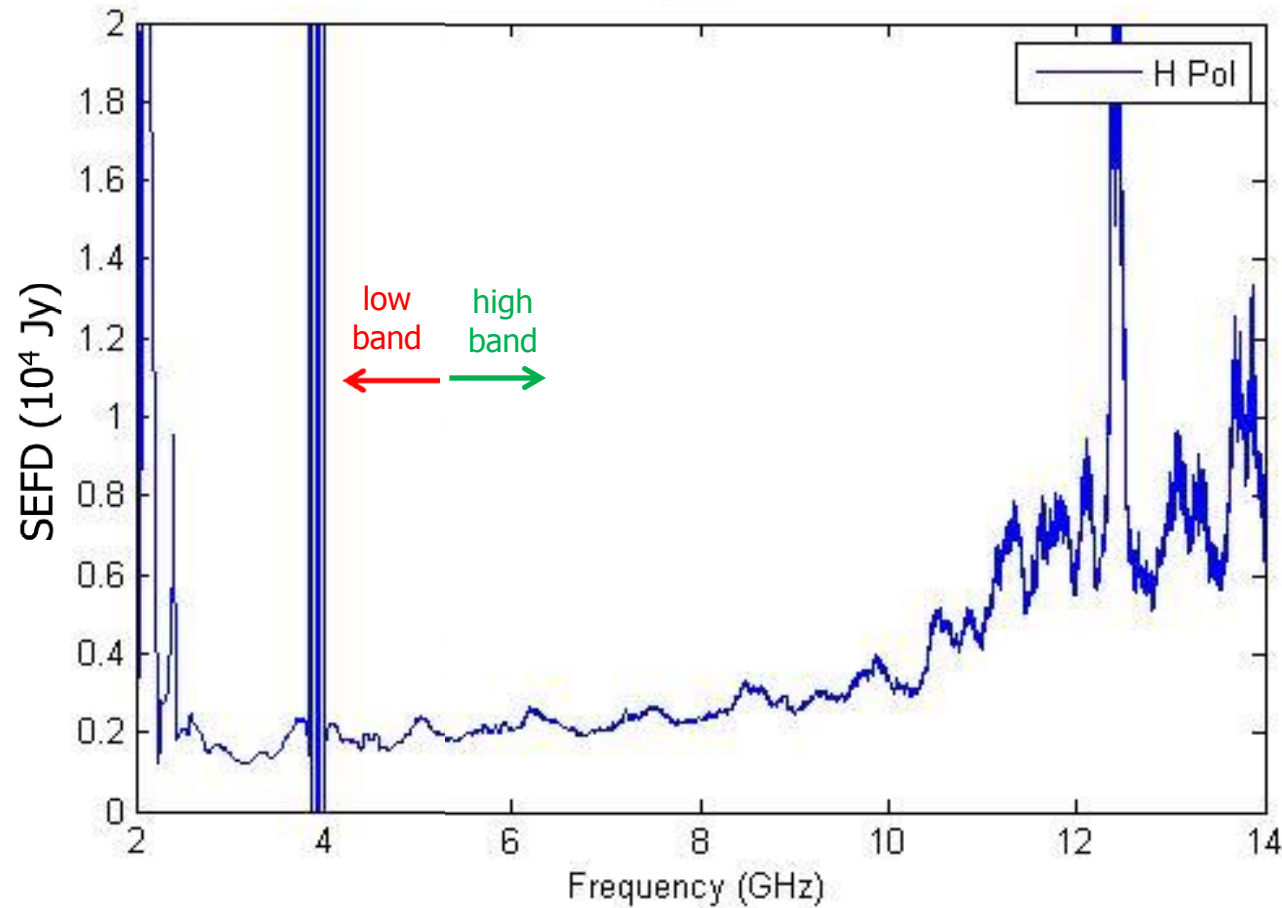


Westford frontend mechanical design

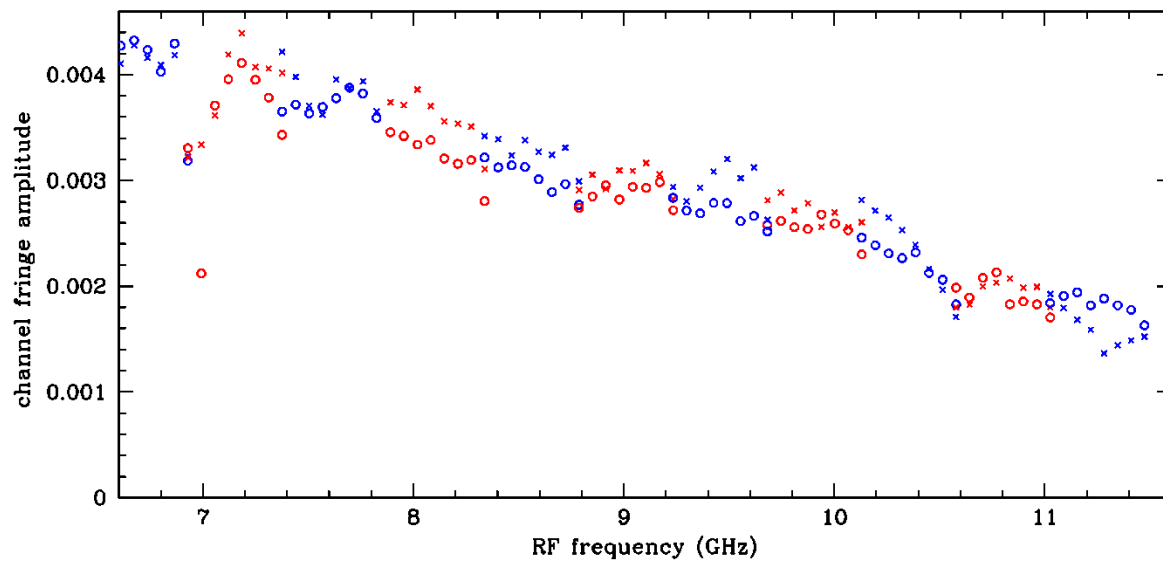
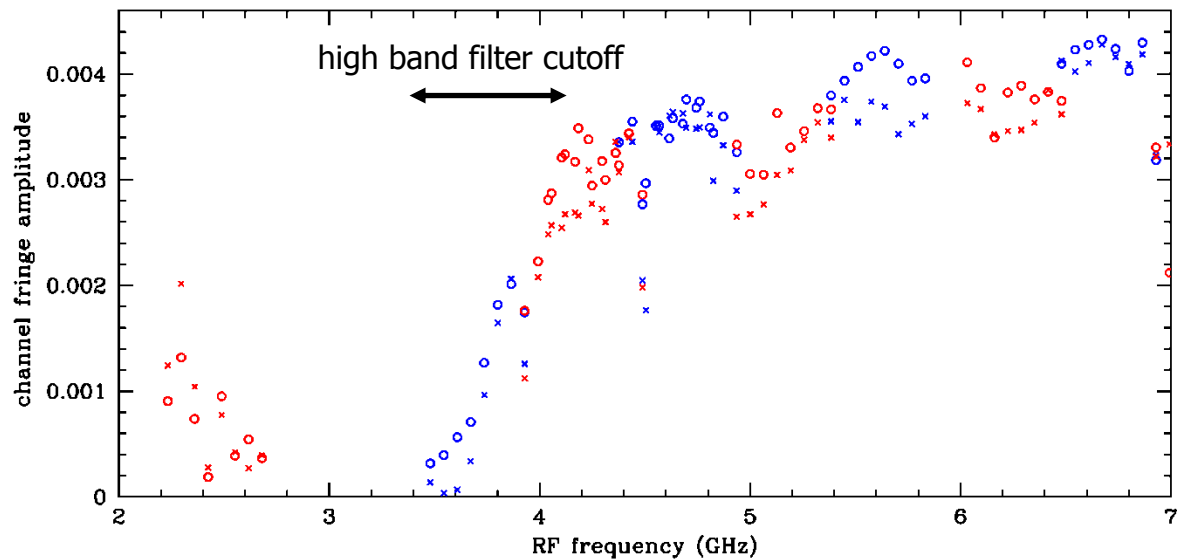


Westford SEFD

Preliminary result –
feed position not yet optimized



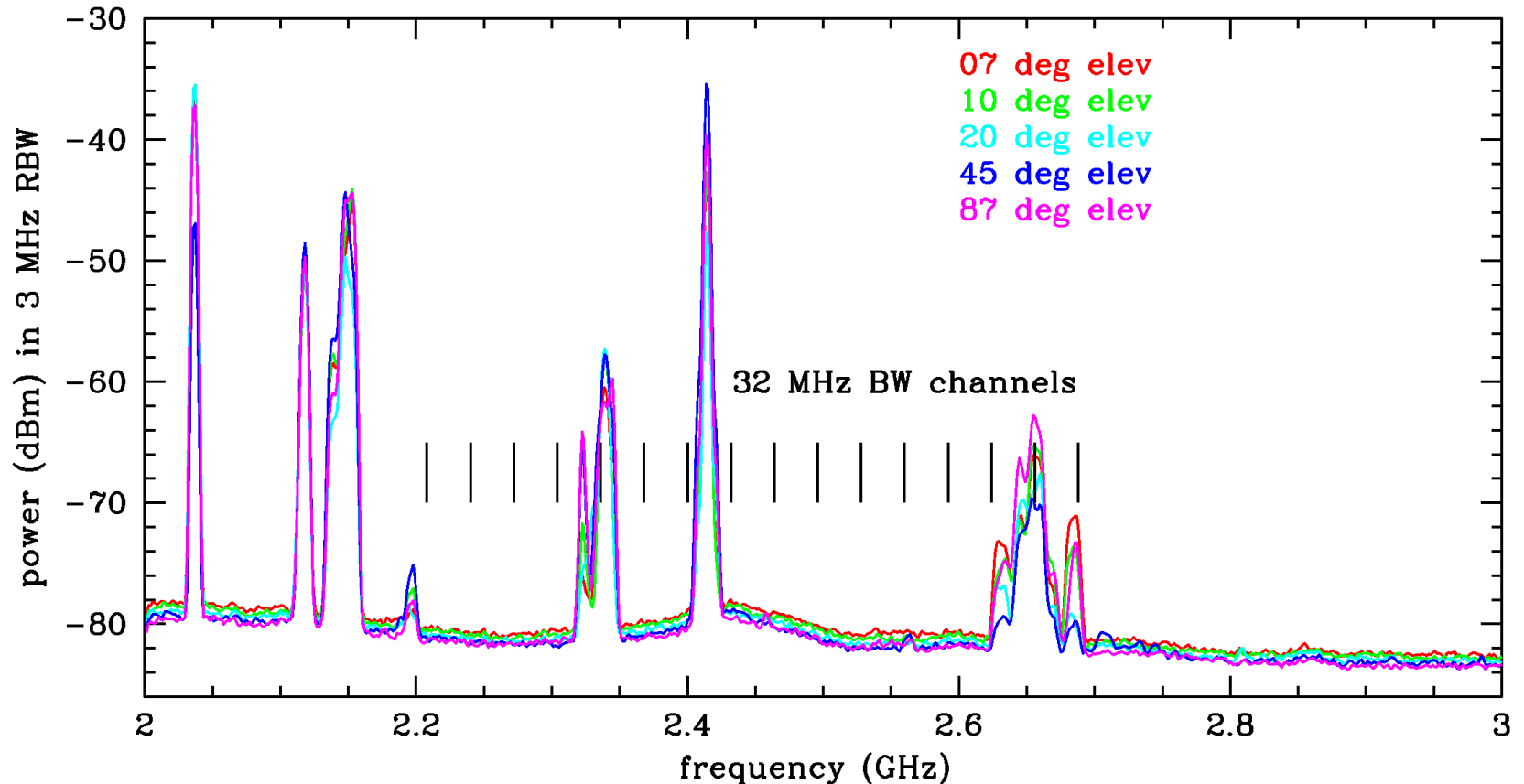
GGAO-Westford fringe amplitude spectrum on 4C39.25



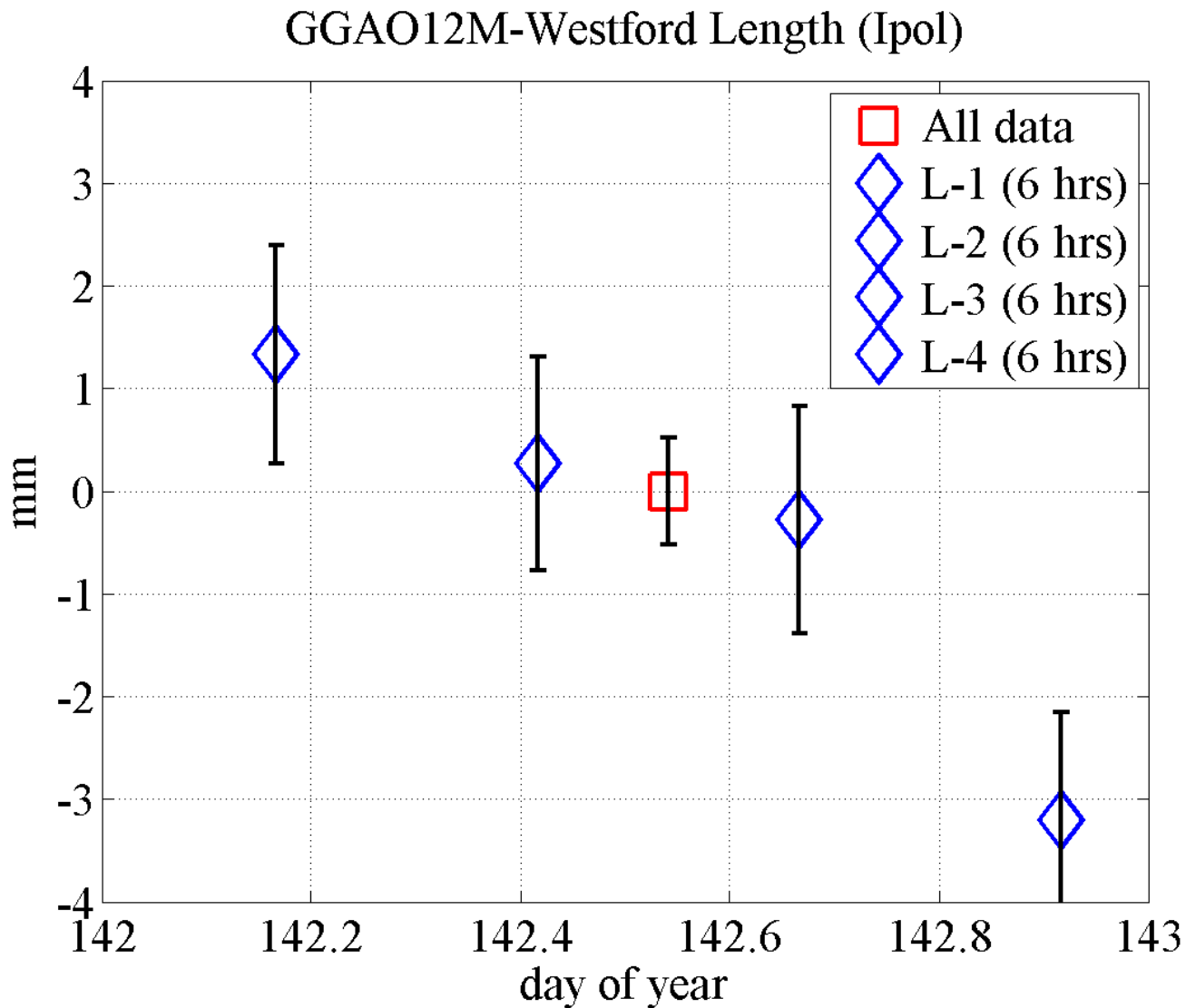
S-band RFI at GGAO

There are enough 32-MHz-wide holes to do geodesy!

GGAO 12m V-pol S-band spectra at 300 deg azimuth

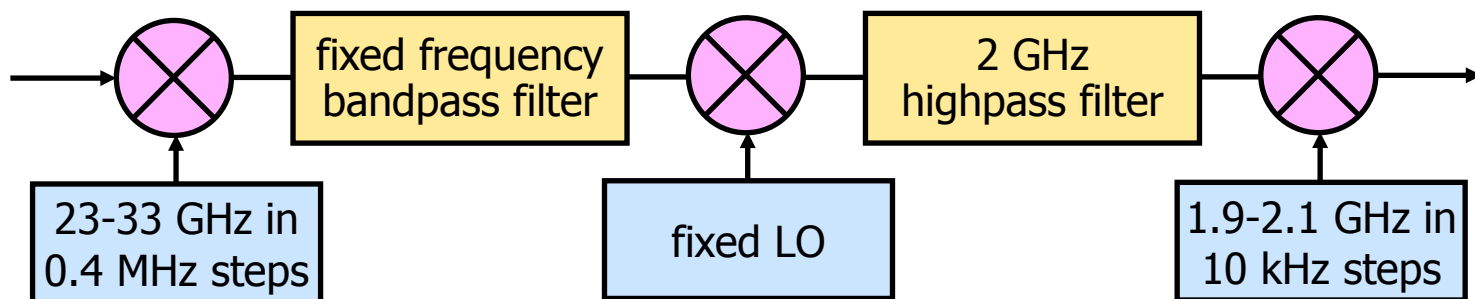


Baseline length repeatability during 24-hour session



Future enhancements

- Field System support for up/down converters, RDBEs, and Mark 6 coming soon!
- Up/down converter
 - With current 0.5-GHz-BW NZ2 filters, present design covers only 1-11.5 GHz input range.
 - Adding a 3rd mixer stage and changing frequencies of bandpass filter and 2nd LO allow:
 - Input range 2-14 GHz to be covered with 0.5-GHz-BW NZ2 or 1.0-GHz-BW NZ1 filters
 - Finer tuning of net LO frequency



Future enhancements -- continued

- Cable Delay Measurement System with sub-ps accuracy under development
 - Needed to measure cable delay variations for 5 MHz to pulse cal generator
 - Thermally driven variations
 - Systematic variations from cable stress in az/el wraps

