

Networks and Processing for LOFAR

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2nd e-VLBI workshop, Dwingeloo

The LOFAR Project

International collaboration between

- ASTRON
- MIT Haystack Observatory
- Naval Research Lab





Membership applications:
 LOIS consortium (Sweden)
 LOPES project (Germany)





LOFAR Key features

- Major new array for 10-240 MHz range
- 400 km across, fixed dipole receptors
- Breakthrough in sensitivity
- Extreme agility in time/space/frequency
- Large instantaneous sky coverage
- Essentially digital



Scientific Versatility



Tozzi et al. (2000)

 Redshifted HI from the Epoch of Reionization

High-z starbursts

Galaxy clusters and the IGM

Cosmic ray distribution, and airshower radio bursts

Steep spectrum and fossil radio galaxies

Supernova remnants and ISM energy budget

Interstellar recombination lines

Nearby pulsars, ghost nebulae

Extrasolar gas giant planetary radio emission

Stellar flares

Interstellar medium propagation effects

Transients, GRB and LIGO event counterparts, buffering

Solar radio studies

CME detection, mapping by IPS, scattering

Extremely high resolution ionospheric tomography

Magnetospheric radar



Lane et al. (2001)



Courtesy: B. Gaensler

CASS/UCSD 2000/07/16 02

Courtesy: B. Jackson

Cosmic Air Showers



LOFAR Technical Concept



- LOFAR is a large distributed radiotelescope:
 - >100 phased array stations
 - Combined in aperture synthesis array
 - 13,000 small "LF" antennas
 - 13,000 small "HF" tiles
 - >20 Tbit/s datanetwork (if full scientific bandwidth)
 - >40 Tflop supercomputer
 - innovative software systems
 - new calibration approaches
 - full distributed control, VO and Grid integration
 - datamining and visualisation

Key technologies

- LOFAR is realized as an aperture synthesis array using multi-beaming phased array techniques.
- LOFAR
 - uses low-cost antennas and
 - relies on broad-band datalinks and
 - on high performance computers
 - to implement the majority of its functionality in (embedded) software.



A distributed embedded system...



1.8 Gpbs per dual polarization antenna220 Gbps per remote station

Antennas and receiver electronics





060303, RBW:3000, Dwingeloo

time [hours]



(in order not to shock you, could be much more...)



Station Processing Functions



Station Processor





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Transient buffering

- Large cyclical RAM buffer per antenna
 Buffer contains full FOV information
 Upon trigger, can point LOFAR post-facto
 Tradeoff between duration, sensitivity
- Implement subsystem, add more RAM later
 Current cost estimate for full array:
 - 3 × 1012 bits/sec for 4 MHz, 2 pols, 14 bits/sample
 - 100 seconds of buffer \Rightarrow \$600k in RAM chips, 2005 prices



Combining stations into an array





Visibility word size	2 Byte
Visibility word size integrated	4 Byte
integration time	10 sec
integration freq resolution	4 kHz
Observation duration	4 hours
Number of SS Visibilities per RS Beam	4950
Correlator output data rate	38.7 Tbps
Correlator processing power	16.9 Tflops
Integrator output data rate	2 Gbps
Integrator processing power	4.8 Tflops
Storage capacity (raw data)	27.8 TB
Selfcal processing power	



Data Transport Network

- Long-range, modest bandwidths (2 .. 20 Gbit/s/station)
 Short-range, high bandwidths (multi Tbit/s)
 - Both: dedicated network (at least in virtual sense)
- Proposed architecture based on 10Gbps
 - Follows industrial/commercial trends
 - Low-cost, short lead times, high market penetration
- A network is more than technology!
 - Even dedicated p2p networks like the LOFAR WAN

 - ⊕ Integration for dual use, …



Central Processing Facility

Handles:

- Correlation
- Beamforming (virtual core and tied array)
- Calibration and associated tasks
- All Sky Monitor imaging
- Other heavy-duty processing (e.g. pulsars)
- Massive processing power
 - Tens to hundreds of Tera-ops (details still under study)
 - Fortunately most problems are "embarrassingly parallel"
 - $\Phi >$ 500 Tbytes of storage
- Implemented via PC clusters
 - Coprocessor boards provide the muscle
 - Special backbone provides the veins





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Correlation

- This is "X" part of FX architecture
 "F" part was done at the stations
- ~5000 baselines for each spectral channel
- Crude calibration applied
 - Data averaging to 1 sec, 10 kHz
 - Output data rate of order 1Gbyte/sec
 - Data storage for further calibration
- Successive stages of calibration, integration
 - Final product smaller data sets, or images



All Sky Monitor

- Image up to 4 steradians every second

 - Correlate, calibrate and image on the fly
 - Examine images for transient sources
 - Time average images to cover many timescales
- Large computational burden options:
 - ✤ Reduce bandwidth to 256 kHz initially?
 - Cluster and beamform VC antennas?
 - Special-purpose number-cruncher?
- Automatic event detection/notification
 - Requires sophisticated algorithms (e.g. reject RFI)







Application Development Framework



Where does LOFAR end?

Observation Modes

- Implemented through "Virtual Instruments"
 - Synthesis Imaging including SelfCal processing
 - Phase Array Operation (no detailed processing)
 - Transient Processing (insufficiently defined yet!)
 - Hooks and stubs for many more

Dataproducts

- Regular and special purpose
 - Global Sky Model (develops in time)
 - Event Lists
 - Sky Models + Residual Maps for synthesis mode
 - Hooks and stubs for anything else



Remote Operations

- LOFAR is an agile, multibeaming instrument
 - Can do many things at once
 - Efficiency demands multiple simultaneous users
 - Resource sharing via internet is integral to LOFAR concept
- General principle for LOFAR operations
 - Bandwidth > internet capacity, CEP task
 - Bandwidth < internet capacity, Science center task</p>
- Limit: functional independence of LOFAR beams
 Design and cost issue



User Data Products

- Data volumes too large for ~2010 internet
 - Lots of online/nearline processing/analysis
 - VO/Grid model: Transport information, not data
 - Sufficient processing to reduce data to manageable volume, suitable for transport for routine operations

Imaging applications

- Widefield (e.g. survey) imaging process on-site
- A Narrow field (targeted) imaging export of visibilities possible
- Non-imaging applications
 - Φ Needs vary by application
 - Voltage time-series might be exported in some cases
 - Some applications (e.g. pulsars) need CEP power





LOFAR Science Centres

- Responsible for the scientific operation of LOFAR and for the support of the broader scientific community.
- Teams consisting of expert LOFAR-users
 - qualified to assessing the instruments behavior and data
 products
 - committed to obtaining the ultimate quality in calibration and scientific integrity.



LOFAR Development Plan

Major milestones

• 100 antenna array + recorder

• 100 antenna array + digital processing + MAC

- 25% of antenna's, 25% of Central Processor
- 2006: Initial Operating Conditions



Software Development Approach

- Time-boxing in 6 months periods
- Large number of well-defined functional modules
 + high-cohesion, low coupling



